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The fate of rotating point plumes in unstratified environment: from free growth to boundary interactions

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8 Plumes generated from a point buoyant source are relevant to hydrothermal vents in lakes 9 and oceans on and beyond Earth. They play a crucial role in determining heat and material 10 transport and thereby local biospheres. In this study, we investigate the development of 11 rotating point plumes in an unstratified environment using both theory and numerical 12 simulations. We find that, in a sufficiently large domain, point plumes cease to rise beyond a 13 penetration height $h_{\rm f}$, at which buoyancy flux from the heat source is leaked laterally to the 14 ambient fluid. The height $h_{\rm f}$ is found to scale with the rotational length scale

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$$h_{\rm f} \sim L_{\rm rot}^p \equiv \left(\frac{F_0}{f^3}\right)^4$$

where F_0 is the source buoyancy flux and $f = 2\Omega$ is the Coriolis parameter (Ω is the 16 rotation rate). In a limited domain, the plume may reach the top boundary or merge with 17 neighboring plumes. Whether rotational effects dominate depends on how L_{rot}^{p} compares 18 to the height of the domain H and the distance between the plumes L. Four parameter 19 regimes can therefore be identified and are explored here through numerical simulation. Our 20 study advances the understanding of hydrothermal plumes and heat/material transport, with 21 applications ranging from subsurface lakes and the oceans in icy worlds such as snowball 22 Earth, Europa, and Enceladus. 23

24 Key words: plumes, rotating convection

25 **1. Introduction**

Turbulent plumes generated by continuous sources or sinks of buoyancy are ubiquitous in nature. Examples include smoke emitted into the atmosphere from chimneys or volcanic eruptions (e.g., Turner 1973; Baines & Sparks 2005), descending convection in high-latitude

29 oceans or under polynyas caused by surface cooling (e.g., Marshall & Schott 1999; Okada

30 *et al.* 2004), and hot springs in volcanic lakes and oceans both on and beyond Earth (e.g.,

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Lupton *et al.* 1985; Varekamp & Rowe 1997; Goodman & Lenferink 2012). Such plumes
can affect the spread of pollutants, transport and mixing of heat and materials (e.g., nutrition
and biomass), and hence the local environment and biosphere. Understanding the growth and
dynamics of such plumes is therefore crucial.

35 Here, we focus on hydrothermal plumes driven by isolated point sources — also referred to as point plumes — in an unstratified environment. This configuration is less investigated 36 37 in the context of Earth oceanography, because weakly stratified oceans are less common on Earth as the ocean is predominantly heated from above, and the buoyancy source from the 38 atmosphere is usually associated with large spatial scales (Schott & Leaman 1991; Marshall 39 & Schott 1999). For these reasons, most previous studies of plume convection have focused on 40 penetrative plume in a stratified ocean or convection driven by wide-spread buoyancy sources 41 42 (see, e.g., Jones & Marshall 1993; Maxworthy 1997; Marshall & Schott 1999). However, the subsurface oceans of icy worlds, for example, snowball Earth (Yang et al. 2012; Ashkenazy 43 44 et al. 2013), Europa (the second moon of Jupiter; Goodman et al. 2004), and Enceladus (the second moon of Saturn; Kang et al. 2022a,b), are likely not stratified or even convectively 45 unstable due to heat sources on the seafloor. In addition, hydrothermal vent activities have 46 been found to exist on Earth (Thomson et al. 1992) and have been hypothesized for icy 47 satellites (Vance & Goodman 2009; Choblet et al. 2017), motivating us to investigate the 48 behavior of rotating point plumes in an unstratified fluid (ocean). 49

Point plumes in an unstratified fluid in the absence of background rotation were first 50 investigated in the milestone study of Morton et al. (1956), hereafter referred to as MTT. 51 By hypothesizing that entrainment varies in proportion to the plume velocity, MTT found 52 analytical solutions for the plume characteristics (width, buoyancy, and speed). Such plumes 53 are found to rise in a conical shape with decreasing buoyancy and speed, and have been 54 studied in the laboratory (e.g., Morton et al. 1956; Boubnov & van Heijst 1994; Ferrero 55 et al. 2022) and numerical experiments (e.g., Abdalla et al. 2009; Bhaganagar & Bhimireddy 56 57 2020).

In the presence of rotation, a point plume would start life as a nonrotating MTT plume, 58 59 and then transit to a cylindrical plume with constant buoyancy anomaly surrounded by rim currents. The transition height and the width of the plume follow the rotational length scale 60 $L_{\text{rot}}^p \equiv (F_0/f^3)^{1/4}$, the only length scale in the system independent of diffusivity or viscosity. Here, F_0 is the buoyancy at the source, and f is the Coriolis parameter. This is supported by 61 62 63 laboratory experiments (Fernando et al. 1998; Goodman et al. 2004; Ferrero et al. 2022) and numerical simulations (Goodman & Lenferink 2012). Upon impact of an upper boundary, 64 cylindrical plumes are found to undergo baroclinic instability, scattering buoyancy away from 65 the source (Fernando et al. 1998; Goodman et al. 2004). 66

However, a rotating plume does not necessarily reach the top surface in an unstratified environment. Water tank experiments by Goodman *et al.* (2004) and oceanic numerical simulations by Kang *et al.* (2022*a*) and Bire *et al.* (2023) demonstrate that concentrated plumes can break up into eddies in the mid-tank/ocean, indicating that the rotation-dominated cylindrical plume may not be the final stage of their evolution. This final stage has rarely been investigated, despite its critical role in determining whether heat and tracers remain concentrated during transport to the surface.

In this paper, we investigate the development of rotating point plumes in an unstratified fluid using both theoretical analysis and numerical simulations. We aim to address two key questions: (1) What factors determine whether a plume remains concentrated as it approaches the top surface or instead disperses before doing so? (2) How does a plume evolve after impacting boundaries? Our paper is organized as follows. A brief description of our numerical method is provided in § 2. The development of a free point plume unconstrained by boundaries is presented in § 3, followed by the development of the plume in a finite domain 81 (§ 4). The effect of source size is discussed in § 5, and key conclusions are summarized in
82 § 6.

83 2. Numerical method

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2.1. Key parameters

The plume is driven by a buoyancy source at the bottom and grows in a rectangular domain that rotates along the *z*-axis at rate Ω . Thus, our system has five free parameters (*L*, *H*, *r*_s, *F*₀, *f*), where *L* and *H* are the width and height of the domain, *r*_s is the radius of the source as the "point" source is actually a finite circular area, *F*₀ (with units of m⁴ s⁻³) is the buoyancy flux imposed at the source, and *f* = 2 Ω is the Coriolis parameter.

The dynamics of the rotating point plumes, where the source size r_s is small enough to be approximated as a geometric point, are primarily governed by F_0 and f. These two physical parameters define a length scale,

$$L_{\rm rot}^p = \left(\frac{F_0}{f^3}\right)^{\frac{1}{4}},\tag{2.1}$$

where the superscript 'p' indicates a point plume. This rotational length scale has been extensively used to describe the characteristic scale of convective plumes in a rotating regime, such as Speer & Marshall (1995), Fernando *et al.* (1998), and Bire *et al.* (2023).

Naturally, the system can be nondimensionlized using L_{rot}^p , yielding the dimensionless parameters

$$\left(\widehat{L},\,\widehat{H},\,\widehat{r}_{\rm s}\right) \equiv \left(\frac{L}{L_{\rm rot}^{p}},\,\frac{H}{L_{\rm rot}^{p}},\,\frac{r_{\rm s}}{L_{\rm rot}^{p}}\right).\tag{2.2}$$

Hereafter, $\widehat{(\cdot)}$ denotes dimensionless parameters. The point source corresponds to $\widehat{r}_{s} \ll 1$, so its geometry is negligible. Instead of using \widehat{L} and \widehat{H} , the following two dimensionless numbers are more commonly used and more convenient for analyses:

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$$(\Gamma, Ro^{*p}) \equiv \left(\frac{L}{H}, \frac{L_{\text{rot}}^{p}}{H}\right), \qquad (2.3)$$

where Γ is the domain aspect ratio, and Ro^{*p} is the natural Rossby number for rotating point plumes that describes the extent to which rotation dominates the dynamics when the plume reaches the surface at height *H* (Goodman *et al.* 2004; Goodman & Lenferink 2012; Bire *et al.* 2023).

When the source size is non-negligible, i.e. $\hat{r}_s \gtrsim 1$, the buoyancy flux per unit area should be the key parameter determining plume dynamics. In this circumstance, the rotational length scale can be formulated as

111 $L_{\rm rot}^a = \left(\frac{F_0}{r_{\circ}^2 f^3}\right)^{\frac{1}{2}},$ (2.4)

where the superscript 'a' represents per unit area. This L_{rot}^{a} has been commonly used in studies of convection driven by wide and homogeneous buoyancy sources (e.g., Stone 1968;

a geometric point, and the influence of the source radius on the plume dynamics is explored

117 further in § 5.

¹¹⁴ Fernando et al. 1991; Bire et al. 2022).

In § 3 and § 4, we focus on the former case, where the buoyancy source can be treated as

2.2. Experimental design

We perform high-resolution simulations for rotating hydrothermal plumes using *Oceananigans.jl*, a GPU-based general circulation model (GCM) written in Julia language with extraordinary computational performance (Ramadhan *et al.* 2020). Our domain spans from -L/2 to L/2 in both x- and y-directions with doubly periodic boundary conditions. In the vertical direction (z-direction), our domain covers z = 0 to z = H. The rotation rate is set to Ω along the z-axis. The initial velocity and buoyancy are both set to zero.

125 At the bottom z = 0, a steady, axis-symmetric, Gaussian-like buoyancy flux is prescribed:

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$$F(x, y) = F_A \exp\left(-\frac{x^2 + y^2}{r_s^2}\right),$$
 (2.5)

where F_A is the maximum buoyancy flux per unit area. The total buoyancy flux is thus given by

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$$F_0 = \iint F(x, y) dx dy = \pi r_s^2 F_A.$$
 (2.6)

In this rotating-plume system, two types of instabilities play a role in the plume devel-130 opment. The first is the Kelvin–Helmholtz instability (hereafter KHI), which occurs during 131 the initial stage, and it triggers the transition from a laminar state to a turbulent state. The 132 second instability is baroclinic instability (hereafter BCI), which occurs after the plume fully 133 develops. It drives lateral mixing between the plume and the ambient fluid. Capturing both 134 instabilities is challenging as BCI occurs at much larger length scales than KHI. To proceed, 135 we set a non-uniform mesh that is refined near the source and gradually coarsens in the upper 136 domain. The dimensionless horizontal and vertical grid sizes, $\widehat{\Delta}_h$ and $\widehat{\Delta}_z$, are given by 137

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$$\widehat{\Delta}_{h} \equiv \frac{\Delta_{h}}{L_{\text{rot}}^{p}} = 0.04, \quad \widehat{\Delta}_{z} \equiv \frac{\Delta_{z}}{L_{\text{rot}}^{p}} = \begin{cases} 0.02 & 1 \leqslant N_{z} \leqslant 300, \\ 0.02 \times 1.01^{(N_{z}-300)} & 301 \leqslant N_{z} \leqslant 450, \\ 0.02 \times 1.01^{150} & 451 \leqslant N_{z} \leqslant 1050, \end{cases}$$

where N_z denotes the vertical grid index, and subscripts 'h' and 'z' refer to horizontal and 139 vertical directions, respectively. The total number of grid cells is 486 in both directions x-140 and y- and is 1050 in the vertical, resulting in a dimensionless domain size of L = 19.44 and 141 $\dot{H} = 66.34$. One experiment using this refined mesh is conducted as the reference case, which 142 is referred to as *Exp-refined* (table 1). The results of this experiment are presented in figure 1. 143 However, due to computational resource limitations, most experiments are performed on a 144 uniform mesh with somewhat lower resolution (the series labeled *Exp-free* and *Exp-res* in 145 table 1). The reduced resolution may delay the onset of KHI due to a compromised ability 146 147 to resolve small-scale turbulence, potentially affecting the subsequent onset of BCI and our 148 main results. To assess the numerical convergence with respect to resolution, a grid sensitivity 149 analysis is presented in \S 3.4.

To monitor the development of plumes, we define the envelope of the plume as the isosurface of $b/b_{\text{max}} = 0.1\%$, where b_{max} is the buoyancy at the source. To diagnose the height of the plume, we use the highest height of the $\overline{b}^y/\overline{b}_{\text{max}}^y = 0.1\%$ isoline, where \overline{b}^y is the buoyancy averaged along the *y*-direction. This approach, equivalent to a side-view perspective, minimizes the need to store large three-dimensional datasets and ensures that the highest plume front is captured regardless of its position.

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156 **3. Development of a free single point plume**

We begin by investigating the development of a single free plume, unaffected by any boundaries. Initially, it behaves like a free-turbulent plume. As rotation becomes significant, it transforms into a Taylor column and eventually becomes destabilized (figure 1).

In this section, we conduct the numerical experiments listed in the upper part of table 1. The results used to visualize the plume evolution are taken from *Exp-refined*. Other experiments were performed on a lower resolution uniform mesh. The sensitivity of our results to grid

resolution is examined in § 3.4.

164 3.1. *Stage-1: free-turbulent plume*

In the first stage, the buoyant fluid emanating from the source is not yet influenced by rotation, so the plume is filled by three-dimensional turbulence. These turbulent motions entrain dense ambient fluid into the plume, reducing its buoyancy and vertical velocity while expanding laterally to form a cone shape (stage-1 in figure 1a).

169 During this stage, F_0 is the only external dimension parameter that determines the plume.

170 As shown in MTT, the radius of the plume R, the upwelling velocity W, and buoyancy

171 anomaly B at the plume center follow,

172
$$R(z) = \frac{6}{5}\alpha z \approx 0.11z,$$
 (3.1)

173
$$W(z) = \frac{5}{6\alpha} \left[\frac{9\alpha}{5} \left(\frac{2F_0}{\pi} \right) \right]^{\frac{1}{3}} z^{-\frac{1}{3}} \approx 4.25 \left(F_0 z^{-1} \right)^{\frac{1}{3}}, \qquad (3.2)$$

174
$$B(z) = \frac{5}{6\alpha} \left[\frac{9\alpha}{5} \left(\frac{2F_0}{\pi} \right) \right]^{-\frac{1}{3}} \left(\frac{2F_0}{\pi} \right) z^{-\frac{5}{3}} \approx 12 \left(F_0^2 z^{-5} \right)^{\frac{1}{3}}, \tag{3.3}$$

where α is the entrainment coefficient (= 0.093 for buoyant plume; Morton *et al.* 1956). Also, we can obtain a volume flux μ and a momentum flux *m* from MTT,

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$$\mu(z) = \frac{6\alpha\pi}{5} \left[\frac{9\alpha}{5} \left(\frac{2F_0}{\pi}\right)\right]^{\frac{1}{3}} z^{\frac{5}{3}} \approx 0.17 \left(F_0 z^5\right)^{\frac{1}{3}}, \qquad (3.4)$$

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$$m(z) = \frac{\pi}{2} \left[\frac{9\alpha}{5} \left(\frac{2F_0}{\pi} \right) \right]^{\frac{2}{3}} z^{\frac{4}{3}} \approx 0.35 (F_0 z^2)^{\frac{2}{3}}, \tag{3.5}$$

which align with the measurements from laboratory experiments (Rouse *et al.* 1952; List &
Imberger 1973; List 1982).

The evolution of the plume in the reference case, *Exp-refine*, is shown in figure 1. Within one rotation period after model initialization, a free-turbulent plume is observed. This plume occupies a clearly defined cone shape (figure 1*b*), which aligns with the one predicted by equation (3.1). Figure 1(*h*) shows the vertical profiles of vertical velocity and buoyancy measured along the center of the plume, W(z) and B(z), both of which decrease as the plume rises and align with equations (3.2) and (3.3).

Note that W(z) and B(z) become singular at z = 0, indicating that MTT is invalid near the source, as the source cannot be treated as a point there. For a circular source with finite area, the point-plume assumption and the MTT model have been shown to be valid only when $z > 24r_s$ (Fernando *et al.* 1998). In our experiment, the invalid region is highlighted in gray in figure 1(*h*). Within this region, our simulation results indeed depart notably from the MTT prediction. Specifically, W(z) is observed to be zero near the bottom due to boundary effects, whereas the theoretical prediction is infinite.

193 effects, whereas the theoretical prediction is infinite.

Exp. name	F_0	f	$L_{\rm rot}^p$	\widehat{L}	\widehat{H}	\widehat{r}_{s}	$\widehat{\Delta}_h$	$\widehat{\Delta}_z$	Г	Ro*p
	$(m^4 s^{-3})$	(s^{-1})	(m)							
Exp-refined	10^{-6}	10^{-6}	1000	19.4	66.3	0.05	0.04	≥0.02	0.293	0.015
Exp-free-1	2.56×10^{-10}	10^{-4}	4	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-2	10^{-8}	10^{-4}	10	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-3	2.56×10^{-6}	10^{-4}	40	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-4	10^{-4}	10^{-4}	100	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-5	2.56×10^{-2}	10^{-4}	400	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-6	1	10^{-4}	1000	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-7	10^{-6}	10^{-6}	1000	50	100	0.05	0.1	0.1	0.5	0.01
Exp-free-8	10^{-6}	10^{-6}	1000	25	100	0.05	0.1	0.1	0.25	0.01
Exp-free-9	10^{-6}	10^{-6}	1000	25	200	0.05	0.1	0.1	0.125	0.005
Exp-res-1	10^{-6}	10^{-6}	1000	20	75	0.05	0.05	0.05	0.267	0.013
Exp-res-2	10^{-6}	10^{-6}	1000	50	100	0.05	0.1	0.2	0.5	0.01
Exp-res-3	10^{-6}	10^{-6}	1000	50	100	0.05	0.1	0.25	0.5	0.01
Exp-res-4	10^{-6}	10^{-6}	1000	50	100	0.05	0.1	0.5	0.5	0.01
Exp-res-5	10^{-6}	10^{-6}	1000	50	200	0.05	0.1	1	0.25	0.005
Exp-wall-1	10 ⁻⁶	10^{-6}	1000	0.1	1	0.005	0.001	0.001	0.1	1
Exp-wall-2	10^{-6}	10^{-6}	1000	0.5	5	0.05	0.01	0.01	0.1	0.2
Exp-wall-3	10^{-6}	10^{-6}	1000	1	10	0.05	0.01	0.01	0.1	0.1
Exp-surf-1	10 ⁻⁶	10^{-6}	1000	1	1	0.005	0.002	0.002	1	1
Exp-surf-2	10^{-6}	10^{-6}	1000	5	5	0.05	0.01	0.01	1	0.2
Exp-surf-3	10^{-6}	10^{-6}	1000	1.5	5	0.05	0.01	0.01	0.3	0.2
Exp-surf-4	10^{-6}	10^{-6}	1000	20	20	0.2	0.04	0.04	1	0.05
Exp-surf-5	9.75×10^{-2}	1.3×10^{-5}	2581	69.7	19.4	0.155	0.139	0.039	3.6	0.052
Exp-surf-6	9.75×10^{-1}	1.3×10^{-5}	4590	39.2	21.8	0.087	0.078	0.044	1.8	0.046

Table 1: Summary of experimental parameters: names of experiments, buoyancy flux F_0 , Coriolis parameter f, rotational length scale L_{rot}^p , dimensionless (hereafter) domain width \hat{L} and height \hat{H} , source radius \hat{r}_s , horizontal and vertical grid sizes $\hat{\Delta}_h$ and $\hat{\Delta}_z$, domain aspect ratio Γ , and natural Rossby number Ro^{*p} . *Exp-refined, Exp-free*, and *Exp-res* experiments are for free plumes discussed in § 3. Other experiments are for plumes constrained by boundaries discussed in § 4. Parameters in *Exp-surf-5* and *Exp-surf-6* are same as experiments shown in figures 2 and 3 in Goodman & Lenferink (2012).



Figure 1: The three development stages of a free rotating point plume in an unstratified environment. (a) Diagram of three stages during plume's growth. (b)–(g) Instantaneous snapshots of the dimensionless buoyancy in the experiment *Exp-refined* in table 1: (b)–(d) side view; (e)–(g) bird's-eye view. Superscripts 'y' and 'z' denote average along y– and z–directions, respectively. (h)–(j) Vertical profiles of normalized vertical velocity (blue) and buoyancy (green) at the plume center. The reference W_{ref} and B_{ref} in panel (h) are calculated from equations (3.2) and (3.3) at $z = 24r_s = 1.2L_{rot}^p$, while in panels (i) and (j), they correspond to w_c and b_c (equations 3.8 and 3.9). In each panel, the theoretical predictions of geometry or vertical profiles (if exist) are represented by dashed curves.

3.2. Stage-2: rotation-dominated cylindrical plume

Since the plume evolves beyond one rotation period, rotation becomes important, and entrained flow from distance is deviated to form a rim current around the plume. This process inhibits lateral material/buoyancy exchange between the plume and the ambient fluid by entrainment, so that the radius of the plume remains almost constant, as can be seen in stage-2 in figure 1(a). The rim current speed is such that it is in thermal-wind balance with the buoyancy contrast inside and outside the plume.

In this stage, the external parameters that determine the plume are F_0 and f, which can be combined into L_{rot}^p (equation 2.1). When the plume is $O(L_{rot}^p)$ away from the source, the Rossby number $Ro \equiv \frac{W}{fR}$ drops below unity, demarcating the transition from the non-rotating to the rotating regime (Ma *et al.* 2020). Following Fernando *et al.* (1998) and Goodman *et al.* (2004), the transitional height h_c and the radius of the cylinder r_c follow

$$h_{\rm c} \approx 6.0 L_{\rm rot}^p, \tag{3.6}$$

$$r_{\rm c} \approx 0.7 L_{\rm rot}^p. \tag{3.7}$$

Within the plume, the vertical velocity and buoyancy are nearly homogeneous and remains constant as the plume rises. They can be estimated from the MTT solution at $z = h_c$,

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$$w_{\rm c} \approx \frac{\mu(h_{\rm c})}{\pi r_{\rm c}^2} \approx 2.2 L_{\rm rot}^p f, \qquad (3.8)$$

$$b_{\rm c} \approx \frac{F_0}{\mu(h_{\rm c})} \approx 0.3 L_{\rm rot}^p f^2.$$
 (3.9)

These scaling laws are found to align with previous lab experiments (Fernando *et al.* 1998; Goodman *et al.* 2004) and numerical experiments (Goodman & Lenferink 2012).

Figure 1(c) shows a snapshot at $t = 9f^{-1}$ in the reference experiment. Below the level of $z = 6L_{rot}^p$, the plume retains a conical shape. However, above this level, the plume is strongly influenced by rotation, taking on a cylindrical shape with a radius of approximately L_{rot}^p . This shape aligns well with the predictions of equations (3.6) and (3.7), as indicated by the black dashed lines. In the cylindrical region, vertical velocity and buoyancy become well mixed both horizontally and vertically, broadly matching the theoretical values given by equations (3.8) and (3.9), as shown in figure 1(*i*).

However, this stage is short-lived for two reasons. First, the plume near the source is quickly deflected from the vertical and begins to precess and oscillate due to the adverse pressure gradient encountered in the rotating environment (Frank *et al.* 2017; Ma *et al.* 2020). This leads to meandering and expansion of the cylindrical plume above. Second, BCI develops, with baroclinic eddies mixing the plume with the surrounding fluid, further driving lateral expansion, as discussed in the following section.

3.3. Stage-3: destabilized plume

The rotation-governed cylindrical plume becomes unstable due to barotropic or baroclinic instabilities (of the rim currents), in line with Kang *et al.* (2022*a*) and Bire *et al.* (2023). After then, the plume will expand horizontally, as eddies scatter the plume fluid into the surrounding environment (stage-3 in figure 1*a*). In our experiments, the destabilization of the cylindrical plume is observed, as the plume spreads horizontally throughout the domain, with its horizontal extent far exceeding r_c (figures 1*d* and 1*g*) and its vertical velocity and buoyancy substantially decreasing (figure 1*j*).

In the experiments listed in the upper part of table 1, the destabilized plumes eventually stop growing in height, despite the absence of stratification. Figure 2(a) shows the evolution

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Figure 2: Evolution of (a) the plume height h_p and (b) the width d_p in the series of experiments labeled *Exp-free* in table 1. The time t is scaled by f^{-1} , and h_p and d_p are scaled by L_{rot}^p . In panel (b), the black dashed line represents the scaling $(ft)^{1/2}$.

of plume height for cases in *Exp-free*, and it is clear that all plumes plateaus at a final height $h_{\rm f}$, $h_{\rm f}$,

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$$h_{\rm f} \approx (55 \pm 5) L_{\rm rot}^p,$$
 (3.10)

which scales with L_{rot}^p . This plateau occurs around $t = 100f^{-1}$, and from then on, the plumes only expand laterally, following $d_p \sim L_{rot}^p (ft)^{1/2}$, where d_p denotes the plume width (figure 2b). This phenomenon has also been observed in a fast-rotating ($Ro^{*p} = 1/60$) tank experiment conducted by Goodman *et al.* (2004) and in numerical simulations for very deep subsurface oceans in icy worlds performed by Kang *et al.* (2022*a*) and Bire *et al.* (2023). The breakdown of concentrated plumes is particularly important for icy worlds, as it dictates whether heat and materials from the vent can be transported to the ice shell without significant dilution.

Here, we propose a physical explanation for the final penetration height $h_{\rm f}$. Inspired by Brickman (1995), Legg *et al.* (1996), and Visbeck *et al.* (1996), we hypothesize that the plume will stop growing when the lateral buoyancy scattering by eddies balances out the buoyancy injection from the source:

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$$\int_0^{h_f} \overline{u'b'} \cdot 2\pi r dz = F_0, \qquad (3.11)$$

where *r* is the radius of the expanded plume, and $\overline{u'b'}$ is the lateral eddy buoyancy flux, with *u'* being the radial eddy velocity and *b'* the buoyancy contrast between the plume and environment. To estimate the eddy buoyancy flux, we need scales for *u'* and *b'*. We assume that *u'* scales with the rim current speed U_{rim} as

$$u' = \gamma U_{\rm rim}, \tag{3.12}$$

where γ is a constant coefficient, in accordance with the theories of Eady (1949), Stone (1972), and Legg *et al.* (1996). The rim current speed U_{rim} can be related to b' through the thermal wind balance,

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$$\frac{\partial U_{\rm rim}}{\partial z} = \frac{1}{f} \frac{\partial b}{\partial r} \implies U_{\rm rim} = \frac{zb'}{fr}.$$
 (3.13)

- For b', we assume that it is well-mixed within the plume, and is diluted from b_c due to the 262 plume expansion (e.g., figures 1d, 1g, and 1j). Thus, it can be scaled as 263
- $b' = CL_{\rm rot}^p f^2,$ (3.14)264

where $C = 0.3r_c^2/r^2$. Here, we have used equation (3.9). Substituting all equations into the 265 buoyancy budget equation (3.11), we obtain 266

$$h_{\rm f} = \frac{1}{\sqrt{\pi\gamma}C} L_{\rm rot}^p. \tag{3.15}$$

From numerical experiments, Legg et al. (1996) estimated $\gamma \approx 0.25$ for plumes in an 268 unstratified environment. Also, in our experiments, the plume width is approximately $10L_{rot}^p$ 269 when it begins to stop ascending (figure 2b), giving $C \approx 0.02$. Substituting these values into 270 equation (3.15) yields 271

 $h_{\rm f} \approx 60 L_{\rm rot}^p$ (3.16)272

273 which quantitatively agrees with our experimental measurements.

3.4. Sensitivity test 274

In our experiments, due to computational limitations, turbulent energy dissipates at the grid 275 scale, which is significantly larger than the Kolmogorov scale. This may influence the growth 276 of turbulent plumes. To assess whether and to what extent the model resolution affects our 277 conclusions, we perform a mesh sensitivity study in this section. We vary $\widehat{\Delta}_h$ and $\widehat{\Delta}_z$ in 278 a series of experiments labeled Exp-res in table 1. This variation effectively modifies the 279 diffusivity and viscosity while maintaining a fixed resolution. 280

Lower resolution delays the onset of KHI and overestimates the height at which the plume 281 transitions to a turbulent state, as illustrates in figure 3. Initially, the plume remains laminar 282 due to the low Reynolds number Re, and then transits to the turbulent state as Re increases. 283 This transition, induced by KHI, has been observed in previous laboratory experiments (e.g., 284 285 Kitamura & Sumita 2011), GCM simulations (e.g., Goodman & Lenferink 2012), and direct numerical simulations (e.g., Ward 2022). The critical height and time for this transition are 286 found to scale with 287

$$L_{\kappa} \equiv \left(\frac{\kappa^3}{F_0}\right)^{\frac{1}{2}}, \quad t_{\kappa} \equiv \frac{\kappa^2}{F_0}, \tag{3.17}$$

where κ is the diffusivity (Ward 2022). In low-resolution experiments, turbulent energy 289 dissipates at larger scales, leading to larger effective κ . This delays the onset of KHI and 290 allows the laminar plume to penetrate to higher heights, potentially exceeding h_c (top left 291 panel in figure 3). Consequently, the MTT stage does not emerge and the plume transits 292 293 directly into a rotationally dominant cylinder (top-right panel in figure 3). As resolution increases, smaller-scale turbulence is better resolved, leading to an earlier onset of KHI and 294 295 the transition to the turbulent plume. This allows the MTT stage to develop and be accurately captured (lower row in figure 3). 296

Furthermore, lower resolution overestimates the final penetration height $h_{\rm f}$, as shown in 297 the central panel of figure 3. This is because the laminar plume in the earlier stage penetrates 298 too high (top-left panel in figure 3). Additionally, a lower-resolution mesh has a reduced 299 300 ability to resolve small-scale baroclinic eddies, leading to an underestimate of lateral eddy buoyancy flux, which also contributes to a higher penetration height. As the resolution 301 increases, $h_{\rm f}$ converges to approximately $55L_{\rm rot}^p$, demonstrating that the mesh adopted before 302 with $\widehat{\Delta}_h = \widehat{\Delta}_z = 0.1$ is sufficient to capture the destabilized plume induced by BCI, while 303

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288



Figure 3: Center: Final penetration height h_f on meshes with different dimensionless (vertical) resolutions. Data points represent the average over the time interval from ft = 100 to ft = 200, with error bars indicating the maximum and minimum values within this interval. Blue line and shading represent $(55 \pm 5)L_{rot}^p$. Upper and lower rows: Side views of the plume shape during its evolution in experiments *Exp-res-4* and *Exp-refined*, respectively. In lower-row panels, gray shading covers the region out of the domain. Several zoomed-in panels are included. Theoretical geometry predictions (if exist) are represented by dashed curves.

minimizing the influence of variations in the onset time of KHI. Consequently, the measured $h_{\rm f}$ (equation 3.10) is robust.

We also adopt different domain extents \hat{L} and \hat{H} in *Exp-refined*, *Exp-free-8*, *Exp-free-9*,

and *Exp-res-1*. The similar plume evolution and h_f shown in figure 2 demonstrate that these

308 domain sizes are sufficient to ensure that the plume is not influenced by boundaries.

309 4. Development of a constrained point plume by boundaries

310 In this section, we investigate the development of point plumes constrained by domain

boundaries, with a particular focus on plume patterns upon encountering the upper boundary.

312 The upper boundary represents either a free surface or an ice shell that covers lakes or oceans.



Figure 4: Parameter diagram for constrained plumes by boundaries (center panel) and simulated plume structures represented by isosurfaces of buoyancy (surrounding panels). The *x*- and *y*-axes are natural Rossby number for point plumes Ro^{*p} and domain aspect ratio Γ , respectively. Circles, squares, triangles, stars, and black dots mark the water tank experiments in Fernando *et al.* (1998) (FCA98) and Goodman *et al.* (2004) (GD+04), and numerical experiments in Goodman & Lenferink (2012) (G&L12), Bire *et al.* (2023) (Bire+23), and this paper (table 1), respectively. Boundaries for four regimes are obtained from equations (3.1), (3.6), (3.7), and (3.10). The parameter regimes for Earth (global-mean $f = 4.4 \times 10^{-5} \text{ s}^{-1}$, gravity $g = 9.8 \text{ m s}^{-2}$, thermal expansivity $\alpha = 3 \times 10^{-4} \text{ K}^{-1}$, heat flux $Q \leq 10$ GW, and ocean depth H = 3 km), Europa (global-mean $f = 1.3 \times 10^{-5} \text{ s}^{-1}$, $g = 1.3 \text{ m s}^{-2}$, $\alpha = 3 \times 10^{-4} \text{ K}^{-1}$, heat flux $Q \leq 10$ GW, and ocean depth $H = 50 \sim 100$ km), and the south pole of Enceladus ($f = 1 \times 10^{-4} \text{ s}^{-1}$, $g = 0.1 \text{ m s}^{-2}$, $\alpha = 10^{-5} \text{ K}^{-1}$, heat flux $Q \leq 5$ GW, and ocean depth H = 40 km) are represented by pink, light green, and light blue shading, respectively.

- 313 With the horizontally periodic domain, we represent a point-plume array that may interact
- with each other during their development. The domain width thus corresponds to the distance between two point sources.
- Here, the parameters that determine the dynamics of the plume are (L, H, L_{rot}^{p}) , which



Figure 5: Instantaneous x-z cross-sections of dimensionless buoyancy b/b_{max} along y = 0 in the experiment *Exp-wall-3* in table 1. Dashed lines represent the plume geometry predicted by MTT model (equation 3.1).

are combined into Γ and Ro^{*p} (equation 2.3). In the parameter space spanned by these two non-dimensional numbers, we identify the following four regimes (figure 4):

- (1) **Lateral-constrained regime**: The domain is too narrow to fit either the non-rotating conical plume, i.e. $\Gamma < 0.22$ (equation 3.1), or the rotation-dominated cylindrical plume, i.e. $\Gamma < 1.4Ro^{*p}$ (equation 3.7). In this regime, the plume will encounter the lateral boundary and become well mixed, so that it cannot remain concentrated when reaching the top boundary.
- (2) **Top-constrained regime (non-rotating)**: The domain is shallow, so the plume will encounter the top boundary as a cone without being influenced by rotation, i.e. $Ro^{*p} > 1/6$ (equation 3.6). However, it is wide enough to fit this conical plume, i.e. $\Gamma > 0.22$ (equation 3.1). In this regime, the plume remains concentrated upon impacting the upper boundary.
- (3) **Top-constrained regime (rotating)**: The rotation is sufficiently fast to constrain the plume into a cylinder before it reaches the top boundary, i.e. $Ro^{*p} < 1/6$ (equation 3.6). In addition, the domain is wide enough to fit this cylindrical plume, i.e. $\Gamma > 1.4Ro^{*p}$ (equation 3.7). In this regime, the plume remains cylindrical and concentrated upon impacting the upper boundary.
- (4) **Unstable regime**: Strongly rotating plumes whose final penetration depth h_f is smaller than domain depth, i.e. $Ro^{*p} < 1/55$ (equation 3.10), in a wide enough domain to fit the cylindrical plume, i.e. $\Gamma > 1.4Ro^{*p}$ (equation 3.7), will become unstable before encountering the upper surface. In this regime, the pattern of the buoyancy source is not maintained.
- 339

4.1. Lateral-constrained regime

To investigate the plume pattern that will first encounter lateral boundaries (equivalent to interactions with other plumes), we carried out three experiments that satisfy the criterion for the lateral-constrained regime, labeled *Exp-wall* in table 1. The plumes in these experiments exhibit similar behavior (bottom-right panels in figure 4).

Snapshots of b/b_{max} at different time steps in *Exp-wall-3* are illustrated in figure 5. The plume is found to encounter the lateral boundaries at $z \approx 4L_{rot}^p$, consistent with the prediction of MTT. Thereafter, the plume merges with adjacent plumes and becomes horizontally wellmixed. This indicates that fluid above this level only "sees" a uniformly distributed buoyancy flux, regardless of the buoyancy flux pattern at the bottom. Over time, these plumes first fill the lower part of the domain with high-buoyancy fluid, because the plume's upwelling must be balanced by surrounding downwelling. Only after the lower domain becomes well mixed



Figure 6: Instantaneous views of dimensionless buoyancy b/b_{max} (color shading) and horizontal flows (vectors) in the experiment Exp-surf-1 in table 1: (a)–(c) x - z cross-sections along red dotted lines in the lower row; (d)-(f) x - y cross-sections at the top. Theoretical geometry predictions (if exist) are represented by black dashed lines.

does the high-buoyancy fluid begin to fill the upper domain, consistent with the "filling-box" 351 effect described by Baines & Turner (1969) and Turner (1969). 352

4.2. Top-constrained regimes

To investigate the plume pattern that will first encounter the top boundary, we conduct a 354 series of experiments labeled *Exp-surf* in table 1, which include both weakly rotating cases 355 (*Exp-surf-1* to *Exp-surf-3*) and strongly rotating cases (*Exp-surf-4* to *Exp-surf-6*). 356

If $Ro^{*p} > 1/6$, rotation has no effect on the rising plume, so the plume maintains the 357 cone shape predicted by the MTT model (equation 3.1) until it meets the upper boundary, as 358 shown in figures 6(a) and 6(d). After that, the buoyant fluid spreads horizontally along the 359 top boundary, forming an anvil layer (figures 6b and 6e). This behavior is similarly reported 360 in tank experiments by Fernando et al. (1998). Eventually, if the domain is relatively narrow, 361 the plume will fill the plane and be pushed downward under the influence of the "filling-box" 362 effect discussed in § 4.1 (figures 6c and 6f). Otherwise, the effect of rotation will eventually 363 dominate, forming anticyclonic flows around a confined plume (figure not shown here). 364

If $Ro^{*p} < 1/6$, rotation would have already started to modulate the plume before 365 encountering the top boundary. As demonstrated in Goodman et al. (2004), in this case, 366 the plume undergoes the first two stages sketched in figure 1(a), and impacts the upper 367 boundary as a cylindrical plume, generating anticyclonic flows (figures 7a, 7b, and 7e). 368 369 Then, the cylindrical plume expands outward due to mass conservation, evolving from the cylinder into a baroclinic cone (figure 7a, 7c, and 7f). The radius of the cone scales with 370



Figure 7: Development of a rotating plume after encountering the top boundary. (*a*) Diagram of three stages during plume's growth. (*b*)–(*g*) Same as figure 6 but for the experiment *Exp-surf-4* in table 1.

371 $L_{\text{cone}} = \sqrt{b'H}/f \sim L_{\text{rot}}^p/\sqrt{Ro^{*p}}$ (Fernando *et al.* 1998; Goodman *et al.* 2004; Goodman

372 & Lenferink 2012), which is the Rossby deformation radius for a two-layer fluid (Pedlosky

1987). We validated this cone scaling though the results are not shown here. Finally, due to baroclinic instability, the primary cone breaks down into multiple secondary conical vortices

 $rac{1}{2}$ (two vortices in figures 7d and 7g). The numbers of these vortices can be estimated using a

³⁷⁶ "heton" model, as shown by Legg & Marshall (1993) and Legg et al. (1996).

377 5. Effect of finite source size

When the source size exceeds the size of the plume characterized by the rotational length scale L_{rot}^{p} , i.e. $\hat{r}_{s} \gtrsim 1$, the source can no longer be treated as a geometric point, but rather a circular region with finite area. This scenario is analogous to convection driven by a widerange buoyancy source in laboratory experiments (Maxworthy & Narimousa 1994; Brickman



Figure 8: Evolution of the plume height $h_p(a)$ scaled by L_{rot}^p and (b) by L_{rot}^a in experiments with wide-range source. Colored circles mark the onset time at which the increase in h_p significantly slows. (c)–(e) Side views of plume shape during its evolution in the experiment with $\hat{r}_s = 3$ (green curve in panels a and b).

1995; Narimousa 1997) and open ocean convection on Earth (Jones & Marshall 1993; Send
& Marshall 1995; Legg *et al.* 1996; Visbeck *et al.* 1996).

With the domain size fixed at $\widehat{L} = 50$ and $\widehat{H} = 100$ and the resolution set to $\widehat{\Delta}_h = \widehat{\Delta}_z = 0.1$, we conduct experiments with different source radii, $\widehat{r}_s = 2$, 3, 4 and 5. Since the source radius r_s is larger than the columnar plume radius r_c , the buoyant fluid emanating from the source has an almost Gaussian-like envelope (figure 8*c*), following the buoyancy flux prescribed at the bottom. Subsequently, the plume forms a cylindrical shape with a radius close to \widehat{r}_s (figure 8*d*). This cylinder then expands laterally, in a way that is similar to point plumes (figure 8*e*), which eventually cause plume height to plateau at $h_f \sim 60L_{rot}^p$ (figure 8*a*), similar to that for point plumes.

It is also evident from figure 8(a) that although the final states of the plumes forced by 392 heat sources of different sizes are similar to each other, the plume height evolutions during 393 their early stages diverge (to the left of the circles in figure 8a). As the heat source widens, 394 the plume rises at a slower rate initially. This is because a larger \hat{r}_s results in a smaller 395 buoyancy flux per unit area, which now primarily governs plume dynamics rather than the 396 total buoyancy flux. Consequently, when nondimensionalized by L_{rot}^{a} (equation 2.4), the early 397 evolution of the plume height from experiments with different \hat{r}_{s} collapses into one line (to 398 the left of the circles in figure 8b). Although the initial development of the plume diverges 399 from the L_{rot}^{p} scaling, it does not carry over to the final stage, which sets the penetration height 400 $h_{\rm f}$. This is because, in all of our experiments, the cylindrical plume is already wide enough 401

to cover the entire buoyancy source before reaching $h_{\rm f}$ (figure 8*d*). Then, the argument that determines the penetration height of the point plume would also apply here, leading to a similar $h_{\rm f}$.

405 6. Summary and discussion

In this study, we investigate rotating plumes in an unstratified environment that are generated by localized buoyant sources at the bottom. We start by considering a plume freely developing without interfering with lateral or top boundaries, then proceed to account for boundary effects, which allows us to determine under which conditions the pattern of buoyancy source can project onto the upper surface.

A rotating point plume that grows freely generally undergoes three stages. In stage-1, 411 with negligible rotation effect, the plume rises as a cone, accompanied by diluted upward 412 velocity and buoyancy (figures 1a, 1b, 1e, and 1h). During this stage, its geometry and 413 dynamics are well described by the MTT model (equations 3.1–3.3; Morton et al. 1956). In 414 415 stage-2, rotation becomes important and forces the plume into a cylindrical shape, within 416 which both momentum and buoyancy are almost uniform (figures 1a, 1c, 1f, and 1i). The geometry and dynamics of this cylindrical plume follow the scaling laws (3.6)–(3.9) (also see 417 Fernando et al. 1998; Goodman et al. 2004). In stage-3, the plume becomes barotropically or 418 baroclinically unstable, leading to lateral loss of mass and buoyancy through eddy transport, 419 which hinders its upward penetration (figure 1a, 1d, 1g, and 1j). 420

We find that freely developed plumes cease to rise at a height $h_{\rm f}$, as the total buoyancy flux released by the source is completely leaked to the ambient by lateral eddy flux (figure 2). Based on this point, we propose a scaling law

$$h_{\rm f} \approx 55 L_{\rm rot}^p$$
. (6.1)

Here, $L_{\text{rot}}^p = (F_0 f^{-3})^{\frac{1}{4}}$ represents the characteristic scale of a rotating plume that is larger than the source size, meaning the source can be treated as a point. When the source size exceeds the plume scale, the source should be treated as a circular area. The early evolution of the plume is then mainly determined by the buoyancy flux per unit area, which leads to the relevant rotational length scale $L_{\text{rot}}^a = (F_0 f^{-3} r_s^{-2})^{\frac{1}{2}}$. However, the final penetration height h_f is close to that of point plumes, since its lateral extend is as wide as the source, which means that the majority of F_0 is contained within the plume (figure 8).

Taking into account the constraints imposed by the boundaries, the evolution of plumes largely depends on their geometry upon encountering these boundaries. We identify four regimes in the non-dimensional parameter space defined by the domain aspect ratio $\Gamma = L/H$ and the natural Rossby number for point plumes $Ro^{*p} = L_{rot}^p/H$, summarized in figure 4. Only in two top-constrained regimes is the pattern of the bottom buoyancy flux preserved when the upper surface is impacted.

In deep convective regions of the Earth in winter, the heat lost to the atmosphere can 438 reach up to $800 \text{ W} \text{ m}^{-2}$ and persist for tens of days, corresponding to buoyancy flux per unit 439 area ranging from 10^{-8} to 10^{-7} m² s⁻³ (Jones & Marshall 1993; Marshall & Schott 1999). 440 The plumes generated in these regions have horizontal scales ranging from several hundred 441 meters to one kilometer, corresponding to $h_{\rm f}$ greater than 6 km (equation 3.10), double the 442 mean depth of the ocean. As a result, the dense cold water can easily descend to the sea floor, 443 forming deep water masses (Jones & Marshall 1993; Send & Marshall 1995). Furthermore, 444 localized hydrothermal vents exhibit heat fluxes ranging from 1 MW to 10 GW (Thomson 445 et al. 1992), corresponding to total buoyancy fluxes ranging from 10^{-3} to $10 \text{ m}^4 \text{ s}^{-3}$. Plumes 446 generated by such vents typically have scales greater than 100 m and can transport to the 447

- surface if the ocean is unstratified (equation 3.10), as shown in figure 4 where we estimateparameter space for Earth's oceans (pink regions).
- Icy satellites, such as Europa (the second moon of Jupiter) and Enceladus (the second moon of Saturn), have been found to host deep global subsurface oceans and exhibit ongoing
- moon of Saturn), have been found to host deep global subsurface oceans and exhibit ongoing
 geological activities (Anderson *et al.* 1998; Hansen *et al.* 2006; Porco *et al.* 2006; Collins
- 452 Goodman 2007). On Europa, the ocean may be 50–100 km deep (Goodman & Lenferink
- 454 2012), and the heat flux for hydrothermal systems is on the order of 0.1–10 GW (Lowell &
- 455 DuBose 2005; Goodman & Lenferink 2012). On Enceladus, the ocean may be 40 km deep
- 456 (Thomas et al. 2016), and the hydrothermal heat flux may range from <0.1 GW following
- 457 Vance & Goodman (2009) to 5 GW following Choblet et al. (2017). In figure 4, we mark
- the parameter space for potential hydrothermal plumes on these two moons in light green
- (Europa) and light blue (Enceladus). The strongest possible plumes on Europa fall into the (rotating) top-constrained regime, suggesting that plumes are able to encounter the ice shell.
- This aligns with the conclusions from the numerical experiments conducted for Europa by
- 462 Goodman & Lenferink (2012), two of which are repeated (*Exp-surf-5* and *Exp-surf-6*) and
- 463 are shown in the top left of figure 4. As for Enceladus, all scenarios considered fall into an
- 464 unstable regime, suggesting that detecting the pattern of hydrothermal vents from the surface
- 465 may be challenging and that nutrients and biosignatures emanated from the hydrothermal
- 466 vents are likely well-mixed when delivered to the water-ice interface, in line with Kang *et al.*
- 467 (2022*a*) and Bire *et al.* (2023).
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REFERENCES

- ABDALLA, IBRAHIM E., COOK, MALCOLM J. & HUNT, GARY R. 2009 Numerical study of thermal plume
 characteristics and entrainment in an enclosure with a point heat source. *Engineering Applications* of Computational Fluid Mechanics 3 (4), 608–630.
- ANDERSON, J. D., SCHUBERT, G., JACOBSON, R. A., LAU, E. L., MOORE, W. B. & SJOGREN, W. L. 1998 Europa's
 differentiated internal structure: Inferences from four galileo encounters. *Science* 281 (5385), 2019–
 2022.
- Ashkenazy, Yosef, Gildor, Hezi, Losch, Martin, Macdonald, Francis A, Schrag, Daniel P. &
 Tziperman, Eli 2013 Dynamics of a snowball earth ocean. *Nature* 495 (7439), 90 93.
- BAINES, P. G. & SPARKS, R. S. J. 2005 Dynamics of giant volcanic ash clouds from supervolcanic eruptions.
 Geophysical Research Letters 32 (24).
- BAINES, W. D. & TURNER, J. S. 1969 Turbulent buoyant convection from a source in a confined region.
 Journal of Fluid Mechanics 37 (1), 51–80.
- BHAGANAGAR, KIRAN & BHIMIREDDY, SUDHEER R. 2020 Numerical investigation of starting turbulent
 buoyant plumes released in neutral atmosphere. *Journal of Fluid Mechanics* 900, A32.
- 494 Bire, Suyash, Kang, Wanying, Ramadhan, Ali, Campin, Jean-Michel & Marshall, John 2022

- 495 Exploring ocean circulation on icy moons heated from below. *Journal of Geophysical Research:* 496 *Planets* 127 (3), e2021JE007025.
- BIRE, SUYASH, MITTAL, TUSHAR, KANG, WANYING, RAMADHAN, ALI, TUCKMAN, PHILIP J., GERMAN,
 CHRISTOPHER R., THURNHERR, ANDREAS M. & MARSHALL, JOHN 2023 Divergent behavior of
 hydrothermal plumes in fresh versus salty icy ocean worlds. *Journal of Geophysical Research: Planets* 128 (11), e2023JE007740.
- 501 BOUBNOV, B. M. & VAN HEIJST, G. J. F. 1994 Experiments on convection from a horizontal plate with and 502 without background rotation. *Experiments in Fluids* **16** (3), 155–164.
- BRICKMAN, DAVID 1995 Heat flux partitioning in open-ocean convection. *Journal of Physical Oceanography* 25 (11), 2609 2623.
- CHOBLET, GAËL, TOBIE, GABRIEL, SOTIN, CHRISTOPHE, BĚHOUNKOVÁ, MARIE, ČADEK, ONDŘEJ, POSTBERG,
 FRANK & SOUČEK, ONDŘEJ 2017 Powering prolonged hydrothermal activity inside enceladus. *Nature Astronomy* 1 (12), 841–847.
- 508 COLLINS, GEOFFREY C. & GOODMAN, JASON C. 2007 Enceladus' south polar sea. Icarus 189 (1), 72–82.
- 509 EADY, E. T. 1949 Long waves and cyclone waves. *Tellus* 1 (3), 33–52.
- FERNANDO, H. J. S., CHEN, R-R. & AYOTTE, B. A. 1998 Development of A Point Plume in the Presence of
 Background Rotation. *Physics of Fluids* 10 (9), 2369–2383.
- FERNANDO, HARINDRA J. S., CHEN, RUI-RONG & BOYER, DON L. 1991 Effects of rotation on convective
 turbulence. *Journal of Fluid Mechanics* 228, 513–547.
- 514 FERRERO, ENRICO, SALIZZONI, PIETRO, IVE, FEDERICA, MANFRIN, MASSIMILIANO, FORZA, RENATO,
 515 BISIGNANO, ANDREA & MORTARINI, LUCA 2022 Effect of rotation on buoyant plume dynamics.
 516 *Heat and Mass Transfer* 58 (7).
- FRANK, D., LANDEL, J. R., DALZIEL, S. B. & LINDEN, P. F. 2017 Anticyclonic precession of a plume in a
 rotating environment. *Geophysical Research Letters* 44 (18), 9400–9407.
- 519 GOODMAN, JASON C., COLLINS, GEOFFREY C., MARSHALL, JOHN & PIERREHUMBERT, RAYMOND T. 2004
 520 Hydrothermal plume dynamics on europa: Implications for chaos formation. *Journal of Geophysical* 521 *Research: Planets* 109 (E3).
- GOODMAN, JASON C. & LENFERINK, ERIK 2012 Numerical simulations of marine hydrothermal plumes for
 europa and other icy worlds. *Icarus* 221 (2), 970–983.
- HANSEN, CANDICE J, ESPOSITO, L, STEWART, A I F, COLWELL, J, HENDRIX, A, PRYOR, W, SHEMANSKY, D &
 WEST, R 2006 Enceladus' Water Vapor Plume. Science 311 (5766), 1422–1425.
- JONES, HELEN & MARSHALL, JOHN 1993 Convection with rotation in a neutral ocean: A study of open-ocean
 deep convection. *Journal of Physical Oceanography* 23 (6), 1009 1039.
- KANG, WANYING, MARSHALL, JOHN, MITTAL, TUSHAR & BIRE, SUYASH 2022a Ocean dynamics and tracer
 transport over the south pole geysers of enceladus. *Monthly Notices of the Royal Astronomical Society* 517 (3), 3485–3494.
- KANG, WANYING, MITTAL, TUSHAR, BIRE, SUYASH, CAMPIN, JEAN-MICHEL & MARSHALL, JOHN 2022b How
 does salinity shape ocean circulation and ice geometry on enceladus and other icy satellites? *Science Advances* 8 (29), eabm4665.
- KITAMURA, SHOGO & SUMITA, IKURO 2011 Experiments on a turbulent plume: Shape analyses. *Journal of Geophysical Research: Solid Earth* 116 (B3).
- LEGG, SONYA, JONES, HELEN & VISBECK, MARTIN 1996 A heton perspective of baroclinic eddy transfer in
 localized open ocean convection. *Journal of Physical Oceanography* 26 (10), 2251 2266.
- LEGG, SONYA & MARSHALL, JOHN 1993 A heton model of the spreading phase of open-ocean deep
 convection. *Journal of Physical Oceanography* 23 (6), 1040 1056.
- 540 LIST, E. JOHN 1982 Turbulent jets and plumes. Annual Review of Fluid Mechanics 14, 189–212.
- LIST, E. J. & IMBERGER, JÖRG 1973 Turbulent entrainment in buoyant jets and plumes. *Journal of the Hydraulics Division* 99 (9), 1461–1474.
- LOWELL, ROBERT P. & DUBOSE, MYESHA 2005 Hydrothermal systems on europa. *Geophysical Research Letters* 32 (5).
- LUPTON, J. E., DELANEY, J. R., JOHNSON, H. P. & TIVEY, M. K. 1985 Entrainment and vertical transport of
 deep-ocean water by buoyant hydrothermal plumes. *Nature* 316 (6029), 621 623.
- MA, YONGXING, FLYNN, MORRIS R. & SUTHERLAND, BRUCE R. 2020 Plumes in a rotating two-layer stratified
 fluid. *Environmental Fluid Mechanics* 20 (1), 103–122.
- MARSHALL, JOHN & SCHOTT, FRIEDRICH 1999 Open-ocean convection: Observations, theory, and models.
 Reviews of Geophysics 37 (1), 1–64.

- MAXWORTHY, T. 1997 Convection into domains with open boundaries. *Annual Review of Fluid Mechanics* 29 (Volume 29, 1997), 327–371.
- MAXWORTHY, T. & NARIMOUSA, S. 1994 Unsteady, turbulent convection into a homogeneous, rotating
 fluid, with oceanographic applications. *Journal of Physical Oceanography* 24 (5), 865 887.
- MORTON, B. R., TAYLOR, GEOFFREY INGRAM & TURNER, JOHN STEWART 1956 Turbulent gravitational convection from maintained and instantaneous sources. *Proceedings of the Royal Society of London. Series A. Mathematical and Physical Sciences* 234 (1196), 1–23.
- NARIMOUSA, SIAVASH 1997 Dynamics of mesoscale vortices generated by turbulent convection at large
 aspect ratios. *Journal of Geophysical Research: Oceans* 102 (C3), 5615–5624.
- OKADA, NAOSUKE, IKEDA, MOTOYOSHI & MINOBE, SHOSHIRO 2004 Numerical experiments of isolated
 convection under polynya. *Journal of Oceanography* 60 (6), 927 943.
- 562 PEDLOSKY, JOSEPH 1987 Geophysical Fluid Dynamics. Springer New York.
- PORCO, CAROLYN C, HELFENSTEIN, PAUL, THOMAS, PC, INGERSOLL, AP, WISDOM, J, WEST, ROBERT, NEUKUM,
 GERHARD, DENK, TILMANN, WAGNER, ROLAND, ROATSCH, THOMAS & OTHERS 2006 Cassini observes
 the active south pole of enceladus. *science* 311 (5766), 1393–1401.
- RAMADHAN, ALI, WAGNER, GREGORY LECLAIRE, HILL, CHRIS, CAMPIN, JEAN-MICHEL, CHURAVY, VALENTIN,
 BESARD, TIM, SOUZA, ANDRE, EDELMAN, ALAN, FERRARI, RAFFAELE & MARSHALL, JOHN 2020
 Oceananigans.jl: Fast and Friendly Geophysical Fluid Dynamics on GPUs. *Journal of Open Source* Software 5 (53), 2018.
- ROUSE, HUNTER, SHUN YIH, CHIA & HUMPHREYS, H. W. 1952 Gravitational convection from a boundary
 source. *Tellus A: Dynamic Meteorology and Oceanography* 4 (3), 201–210.
- SCHOTT, FRIEDRICH & LEAMAN, KEVIN D. 1991 Observations with moored acoustic doppler current profilers
 in the convection regime in the golfe du lion. *Journal of Physical Oceanography* 21 (4), 558 574.
- SEND, UWE & MARSHALL, JOHN 1995 Integral effects of deep convection. *Journal of Physical Oceanography* 25 (5), 855 872.
- SPEER, K. G. & MARSHALL, J. 1995 The growth of convective plumes at seafloor hot springs. *Journal of Marine Research* pp. 1025–1057.
- STONE, PETER H. 1968 Some properties of hadley regimes on rotating and non-rotating planets. *Journal of Atmospheric Sciences* 25 (4), 644 657.
- STONE, PETER H. 1972 A simplified radiative-dynamical model for the static stability of rotating atmospheres.
 Journal of Atmospheric Sciences 29 (3), 405 418.
- THOMAS, P.C., TAJEDDINE, R., TISCARENO, M.S., BURNS, J.A., JOSEPH, J., LOREDO, T.J., HELFENSTEIN, P. &
 PORCO, C. 2016 Enceladus's measured physical libration requires a global subsurface ocean. *Icarus* 264, 37–47.
- THOMSON, RICHARD E., DELANEY, JOHN R., MCDUFF, RUSSELL E., JANECKY, DAVID R. & MCCLAIN, JAMES S.
 1992 Physical characteristics of the endeavour ridge hydrothermal plume during july 1988. *Earth* and Planetary Science Letters 111 (1), 141–154.
- TURNER, J S 1969 Buoyant plumes and thermals. Annual Review of Fluid Mechanics 1 (Volume 1, 1969),
 29–44.
- 590 TURNER, J. S. 1973 Buoyancy Effects in Fluids. Cambridge University Press.
- VANCE, STEVEN & GOODMAN, JASON 2009 Oceanography of an Ice-Covered Moon. In: Pappalardo, R.T.,
 McKinnon, W.B., Khurana, K.K. (Eds.), Europa. University of Arizona Press.
- VAREKAMP, JOHAN C. & ROWE, GARY L. 1997 Exotic fluids in the exosphere: When hades meets apollo.
 Eos, Transactions American Geophysical Union 78 (23), 237–238.
- VISBECK, MARTIN, MARSHALL, JOHN & JONES, HELEN 1996 Dynamics of isolated convective regions in the
 ocean. Journal of Physical Oceanography 26 (9), 1721 1734.
- WARD, DANIEL ROBERT 2022 Unifying laminar and turbulent dynamics of plumes. Ph.d. thesis, University
 of Leeds.
- YANG, JUN, PELTIER, W. RICHARD & HU, YONGYUN 2012 The initiation of modern "soft snowball" and "hard
 snowball" climates in ccsm3. part ii: Climate dynamic feedbacks. *Journal of Climate* 25 (8), 2737 –
 2754.