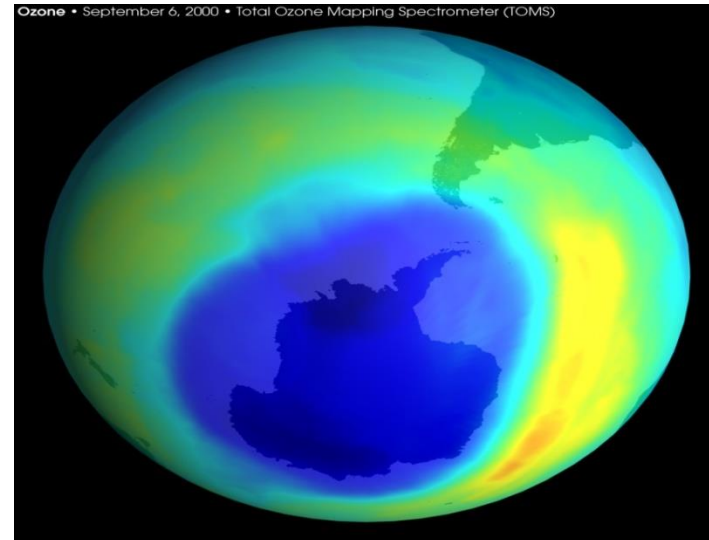


# Large radiative forcing due ozone depletion offset by albedo over Antarctica

G.Chiodo<sup>1</sup>, L.Polvani<sup>1,2</sup> and M.Previdi<sup>2</sup>

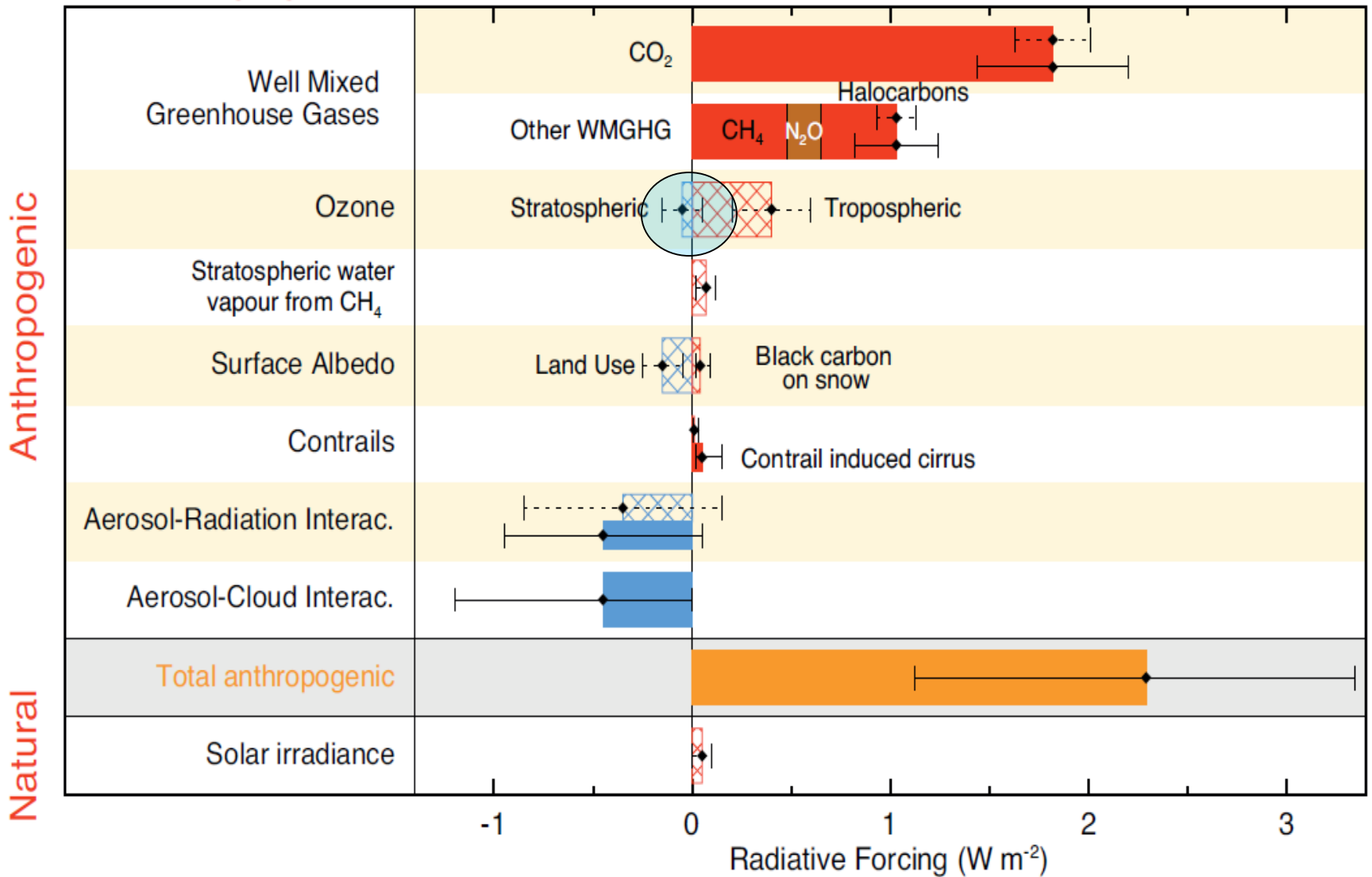
1.Applied Physics and Applied Mathematics, Columbia University, NY

2.Lamont Doherty Observatory, Columbia University, NY

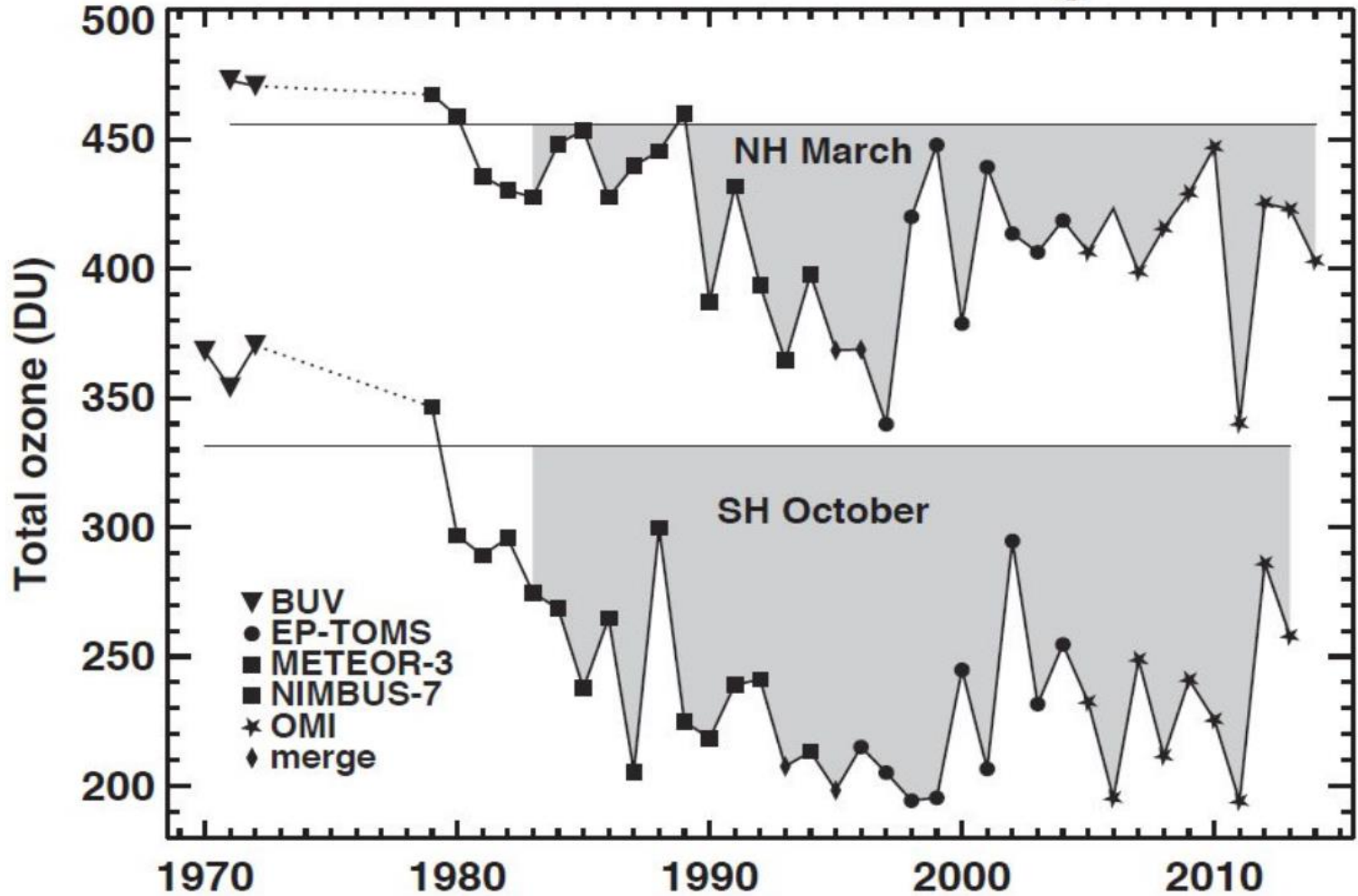


# Radiative forcing of climate between 1750 and 2011

Forcing agent



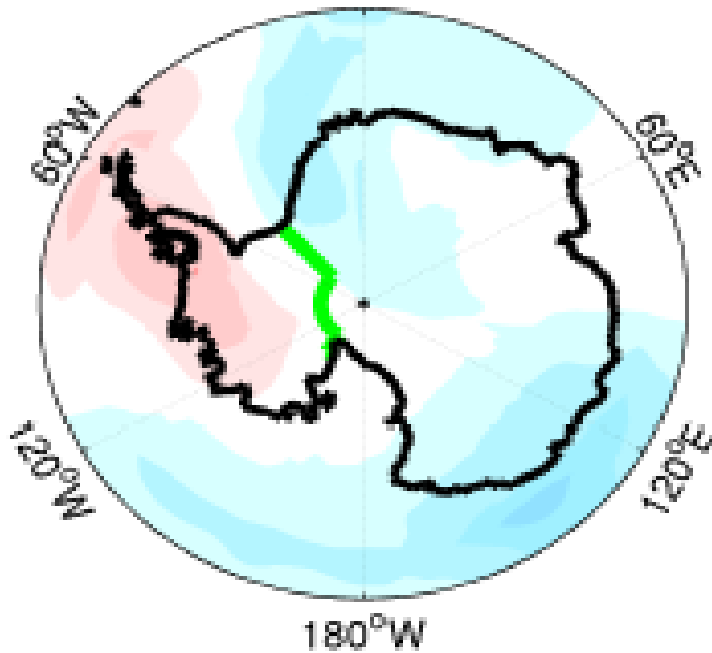
# 63°-90° total ozone average



WMO O3 assessment 2014

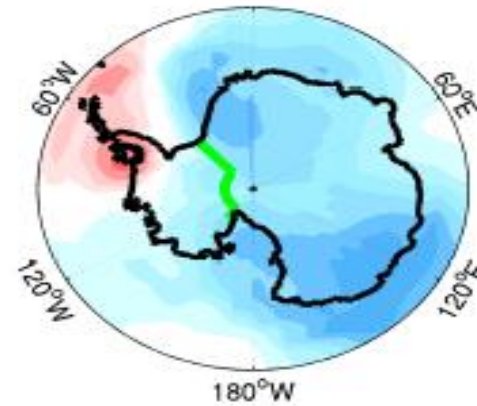
# Observed surface temperature trends

(a) OBS (ANN)

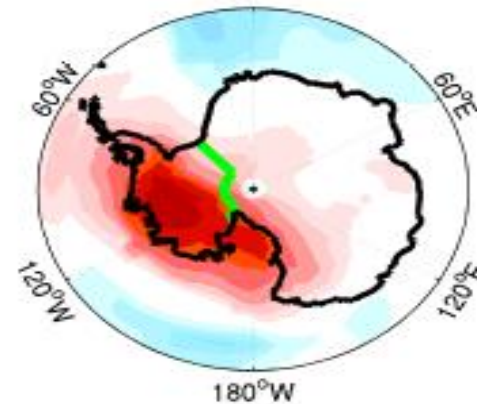


Smith et al., 2016

(d) OBS (MAM)

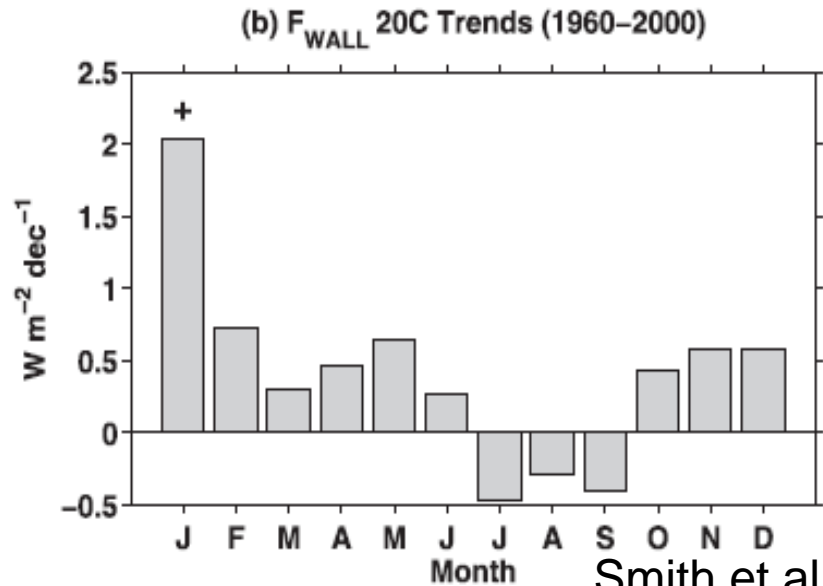
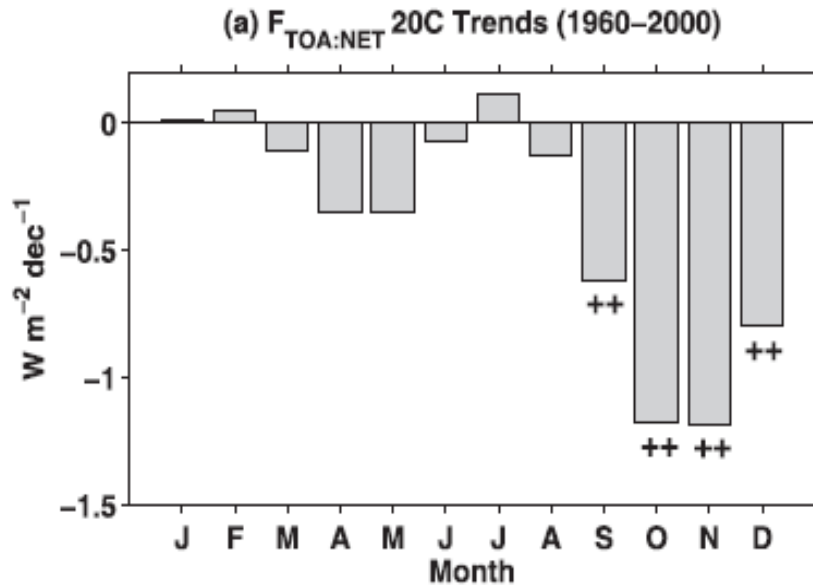


(g) OBS (SON)



Did ozone depletion play a role?

# Effect of ozone depletion on Antarctic (atmospheric) energy budget



Smith et al., 2013

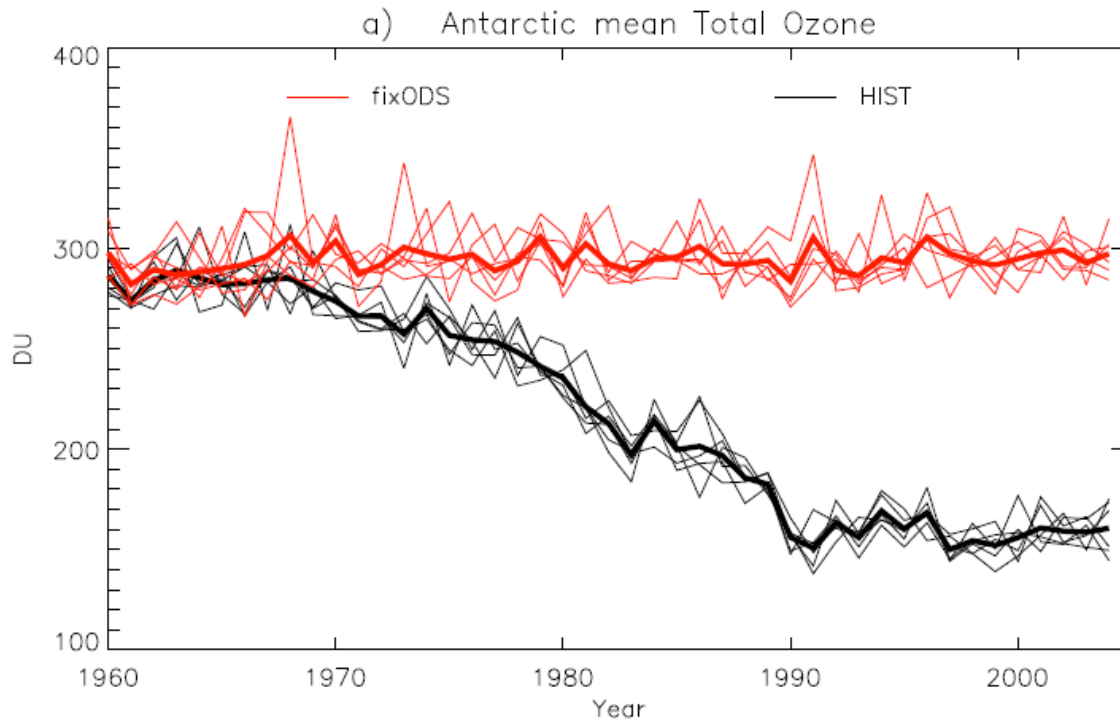
- O<sub>3</sub> depletion leads to energy “deficit” at TOA
- Energy deficit balanced by increased energy flux convergence (right)

**What is the impact of O<sub>3</sub> depletion on surface energy budget?**

# Method

CESM-WACCM4 model, 1.9x2.5 deg, coupled ocean and sea-ice

**6 members with O3 depl (“HIST”) vs 6 members without (“fixedODS”)**  
(both include GHGs)



O3-hole effect : HIST-fixODS (1990-2005) OND season

# Radiative forcing from stratospheric ozone depletion

Offline PORT (Conley et al., GMD 2013)

## Response to perturbation p

$$\begin{aligned}\frac{dT}{dt} &= H(T, C) = Q(T, C) + D(T, C), \\ \frac{dT_p}{dt} &= H(T_p, C_p) = Q(T_p, C_p) + D(T, C).\end{aligned}$$

$$D(T, C) = D(T_p, C_p)$$

## Stratospheric adjustment

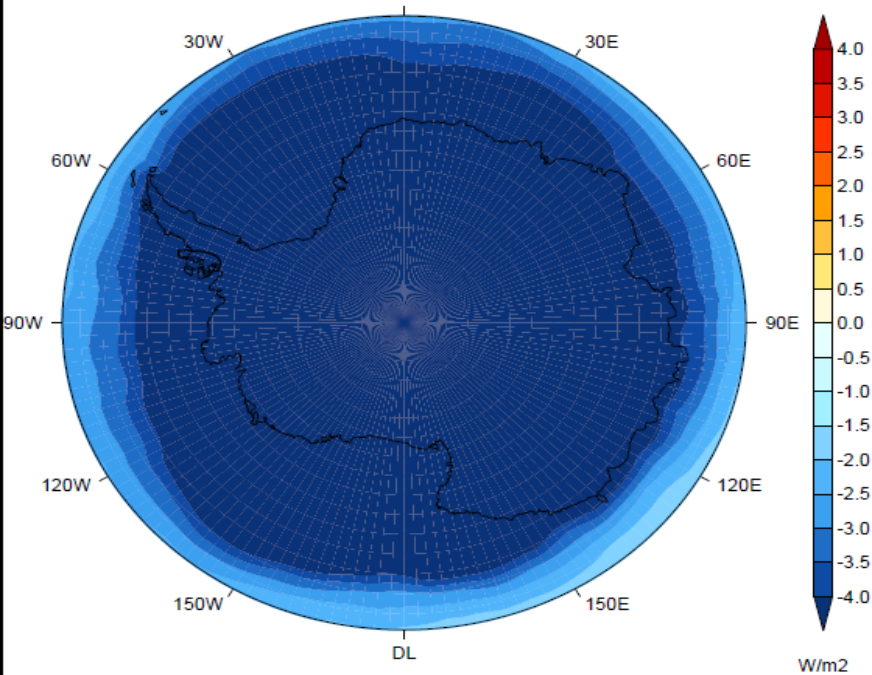
$$\begin{aligned}\frac{dT_{sa}}{dt} &= H(T, C) - H(T_p, C_p) \\ &= Q(T, C) - Q(T_p, C_p).\end{aligned}$$

-2000 vs 1955 ozone from **HIST**

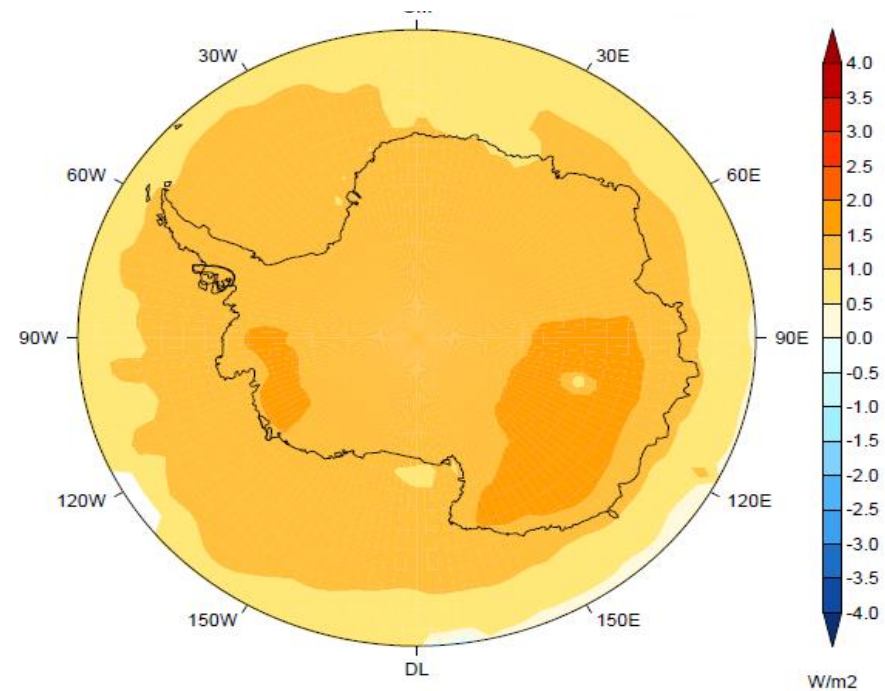
-Same cloud cover, sea-ice, and RAGs (CO<sub>2</sub>,H<sub>2</sub>O)

# Adjusted TOA forcing (OND)

**SW**



**LW**



Net adjusted TOA forcing is 3-4 W/m<sup>2</sup>

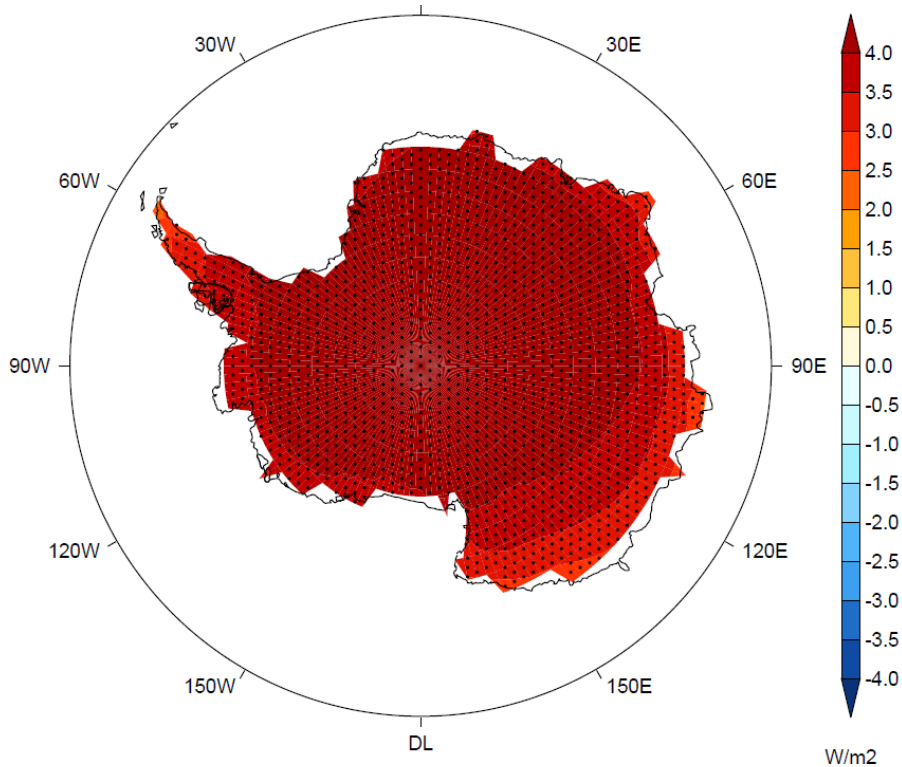
**What is the direct effect of O<sub>3</sub> depletion at the surface?**



# Changes in surface downwelling SW

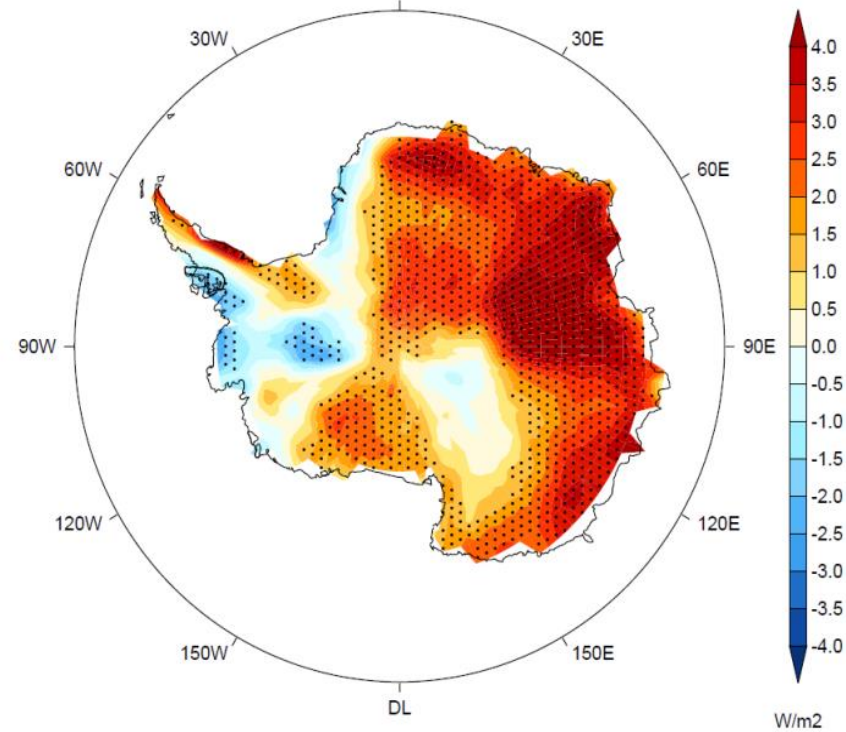
HIST-fixODS (1990-2005)

## clear sky SWd



+ 3.8 W/m<sup>2</sup>

## all sky SWd

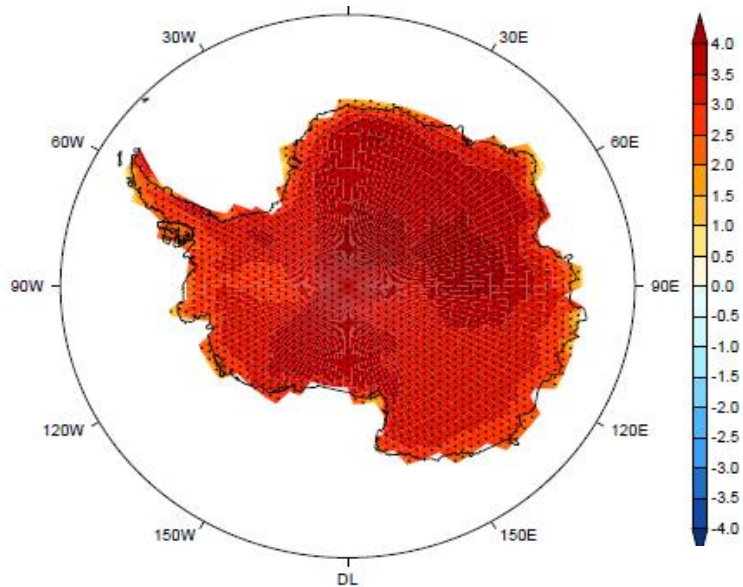


+ 1.7 W/m<sup>2</sup>

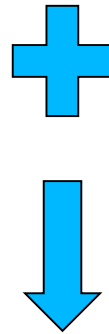
# Partial compensation by clouds (near IR)

HIST-fixODS (1990-2005)

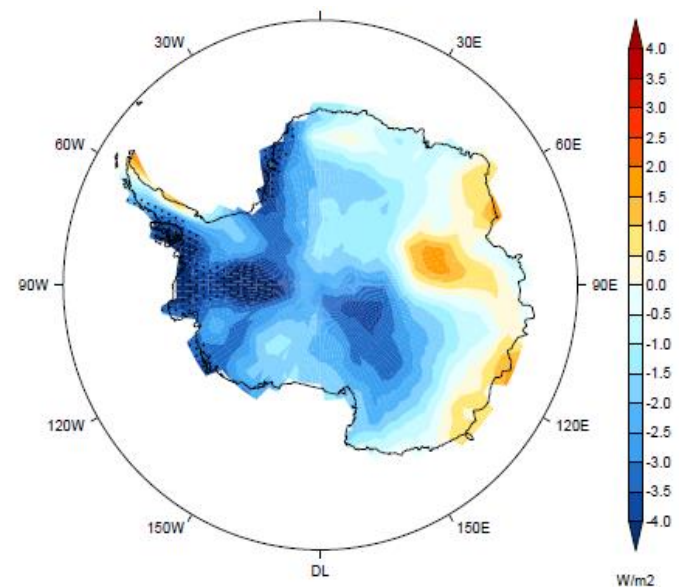
**UV+VIS SWd**



+ 3.4 W/m<sup>2</sup>

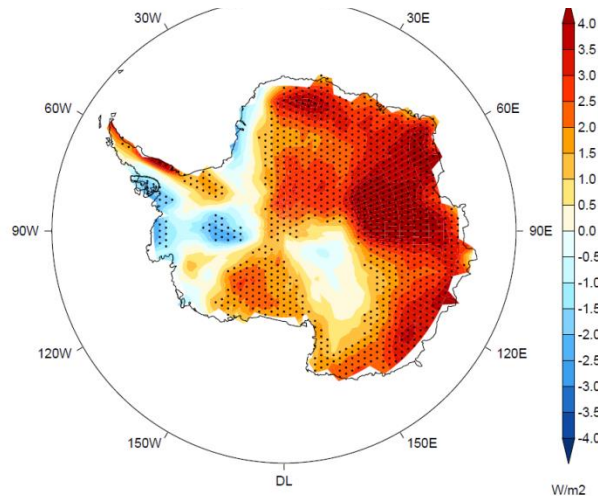


**near IR SWd**



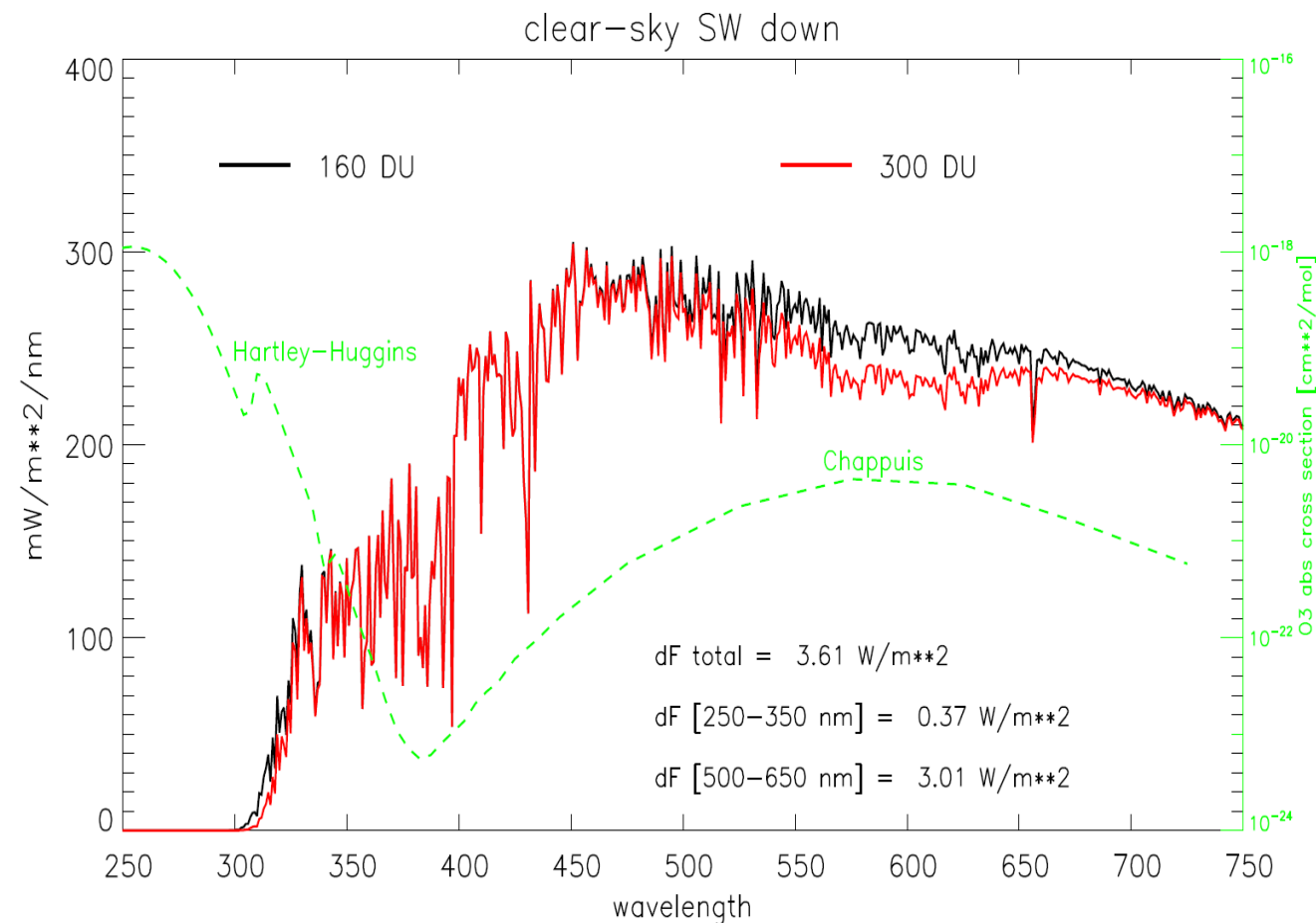
- 1.7 W/m<sup>2</sup>

**all sky SWd**



# Is the large surface SW perturbation realistic?

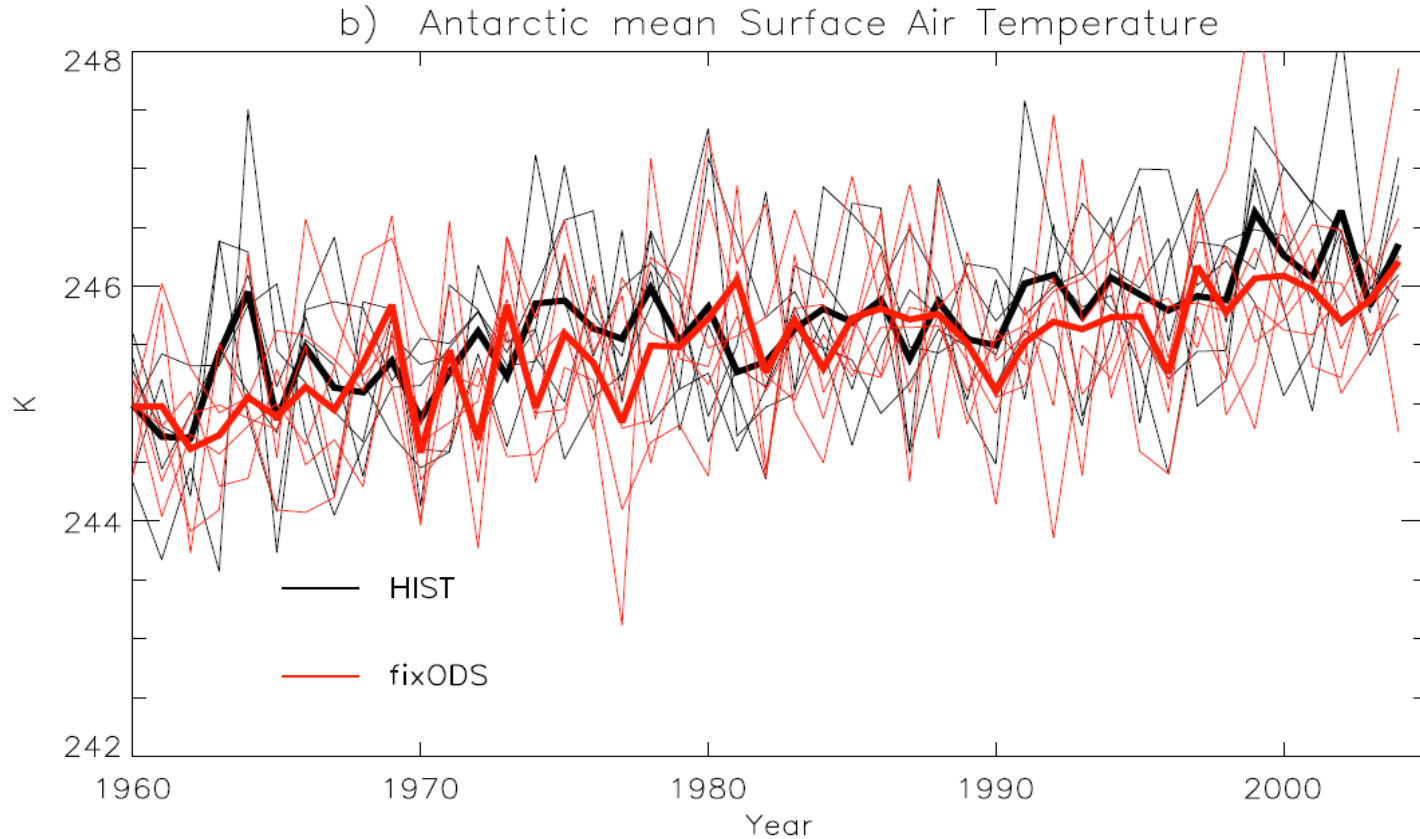
LibRadTran LbL (CCMVal, 2010)  
single column, clear-sky at 80SZA, pseudo-spherical



$dF$  of 3.61 W/m<sup>2</sup>,  
close to CESM-  
WACCM

Most of  $dF$  in 500-650  
nm; Chappuis Band

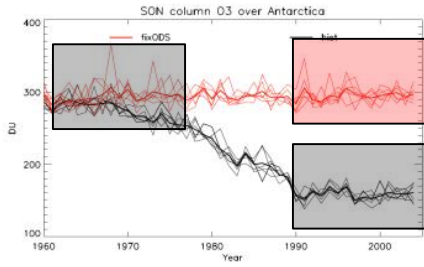
# Does O3 depletion warm Antarctica in OND?



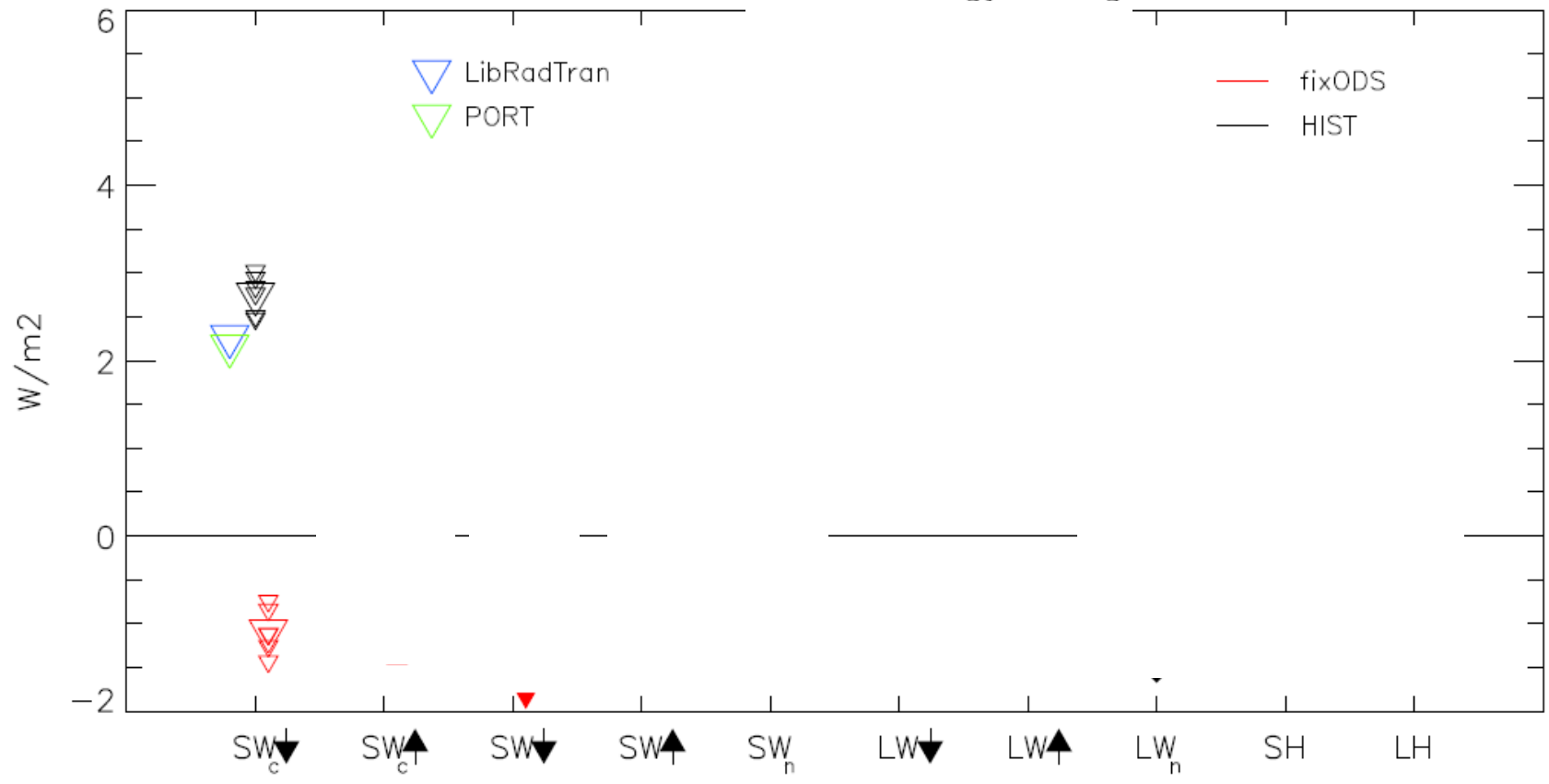
**... NO!**

Very similar warming trend in both HIST and fixODS

# Why does O3 depletion **not** affect Antarctic surface climate?



Antarctic surface energy budget



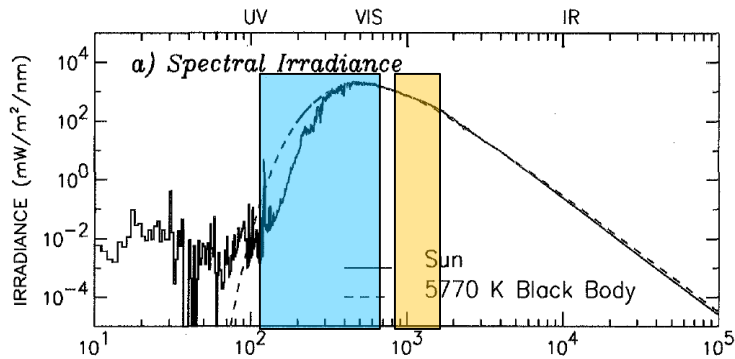
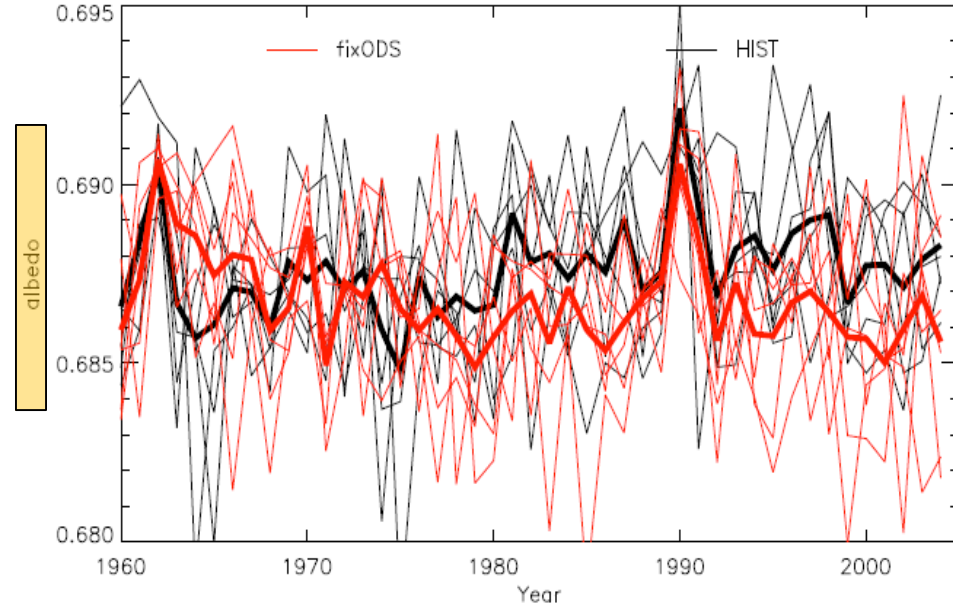
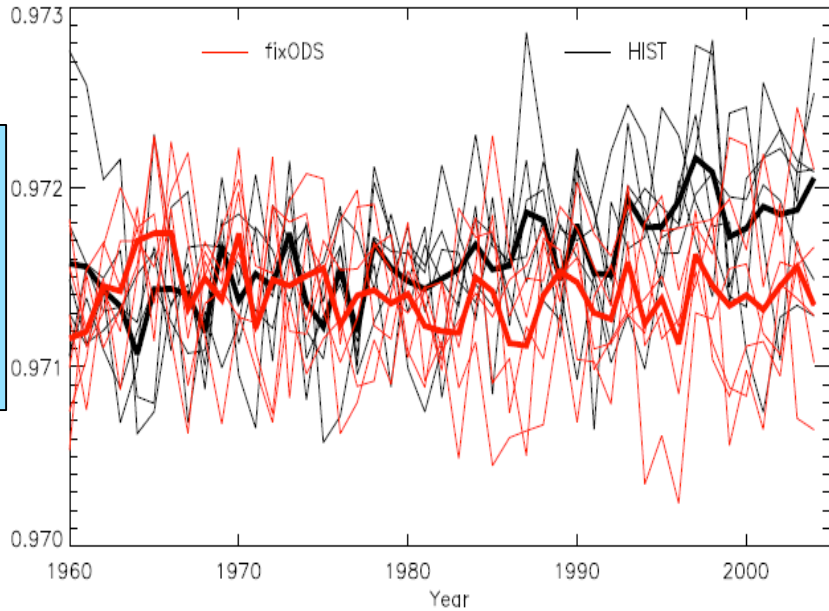
Post minus pre O3 hole period (1990-2005 clim minus 1960-1975 clim)

# Surface albedo

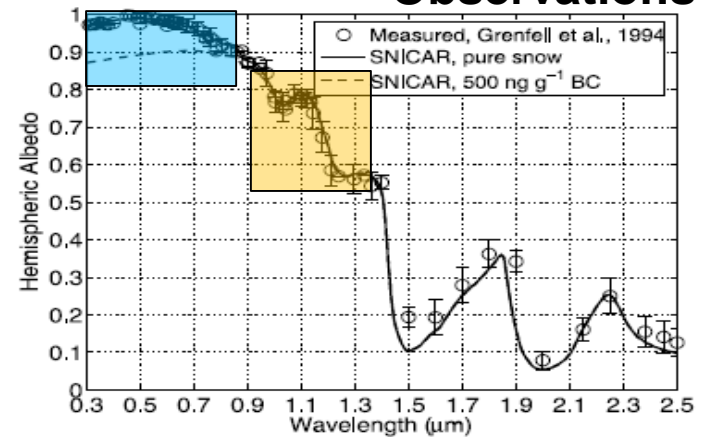
## UV/VIS albedo

## WACCM4

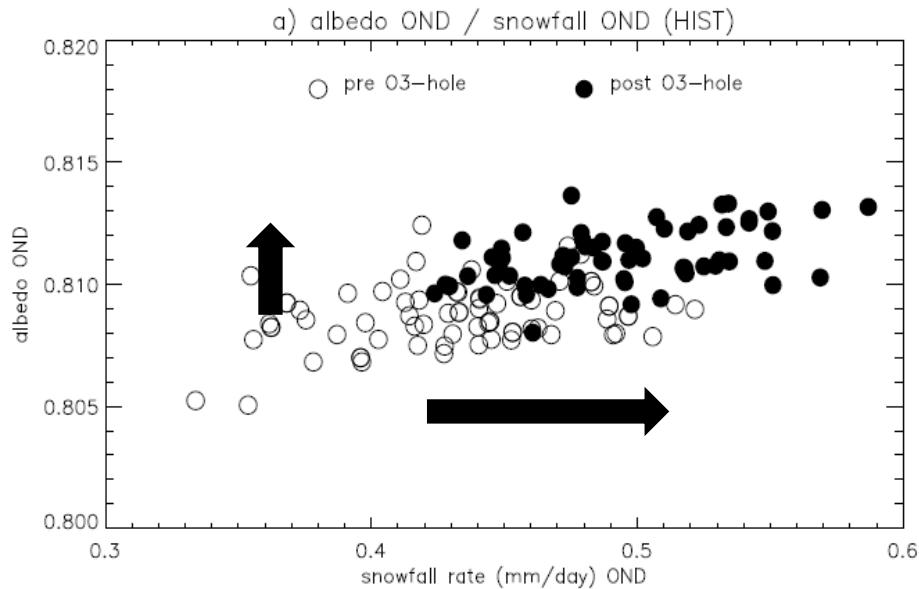
## nearIR albedo



## Observations



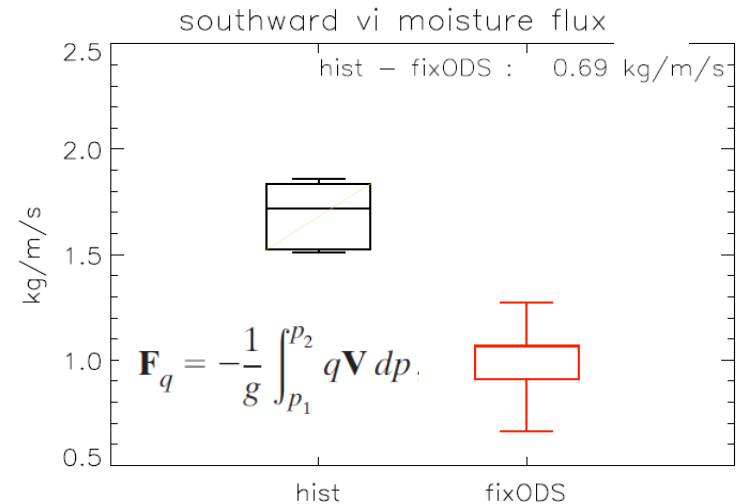
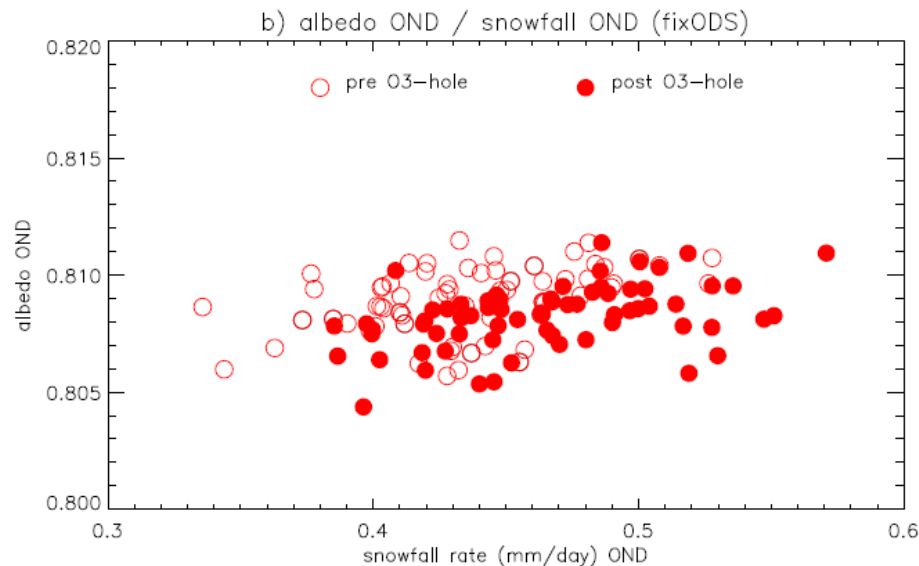
# Cause of UV+VIS albedo increase



Increase in albedo (0.3%) related to ozone-hole induced **increase in Antarctic snowfall**

$$r_e(t) = \left[ r_e(t-1) + dr_{e,dry} + dr_{e,wet} \right] f_{old} + r_{e,0} f_{new} + r_{e,rfz} f_{rfz}$$

Albedo controlled by snowfall rate (snow aging scheme in CLM4)

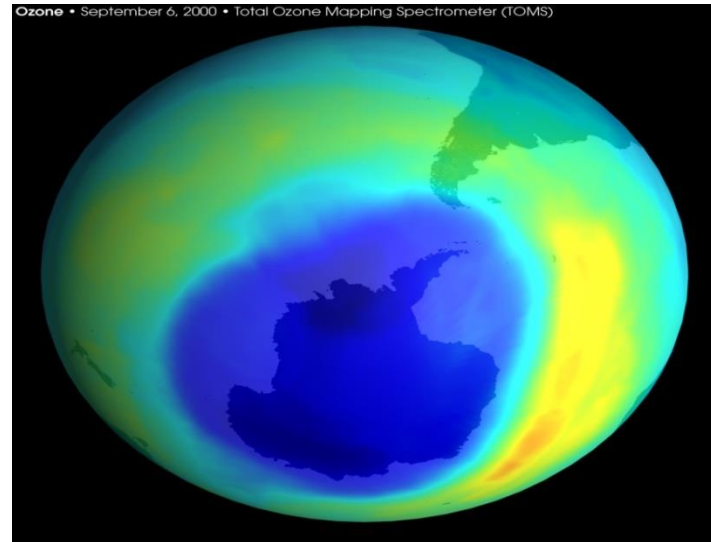


# Conclusions

- 1) O<sub>3</sub> depletion leads to a detectable increase in downwelling SW over Antarctica, consistent with LbL calculations.
- 2) The SW perturbation is largely cancelled by albedo, and to lesser extent, by albedo trends induced by snowfall, which make Antarctica brighter. As a result, the net SW forcing is very small.
- 3) Thus, the ozone hole has a negligible radiative effect on Antarctica, and is thus unlikely to induce continental scale temperature changes.

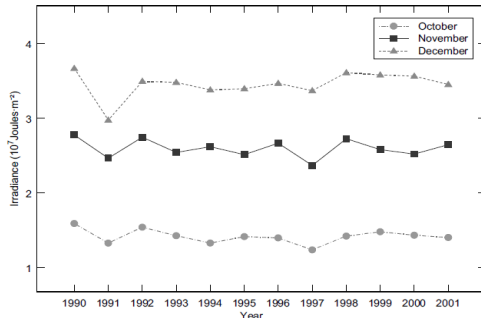


# Thank you

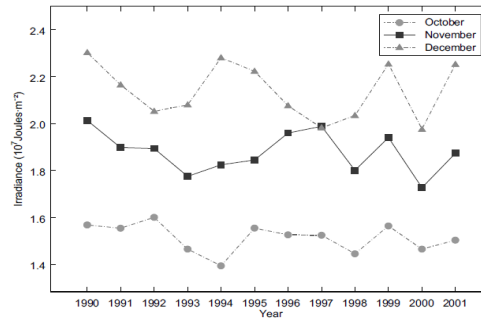


# Did downwelling SW radiation increase in the real world ?

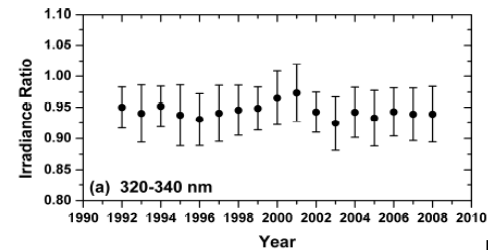
McMurdo



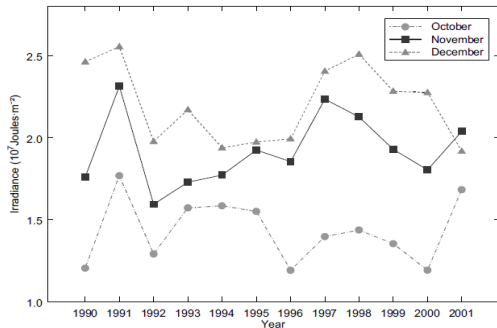
Ushuaia



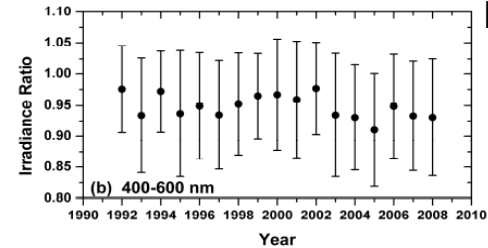
South Pole



Palmer

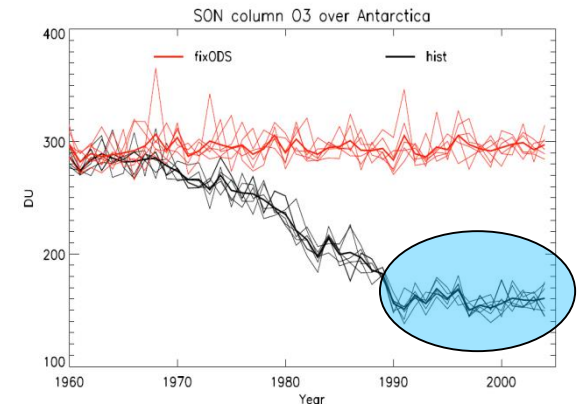


Frederick et al. 2007



Frederick and Hodge acp 2011

... hard to say, since in-situ measurements started in 1990s



# Did Antarctic albedo increase in the real world ?



Available online at [www.sciencedirect.com](http://www.sciencedirect.com)



Remote Sensing of Environment 112 (2008) 646–667

Remote Sensing  
of  
Environment

[www.elsevier.com/locate/rse](http://www.elsevier.com/locate/rse)

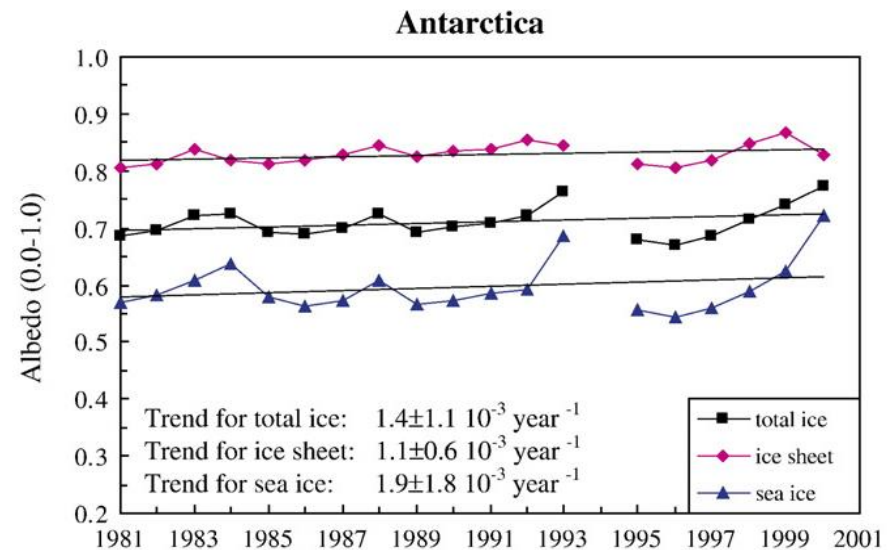
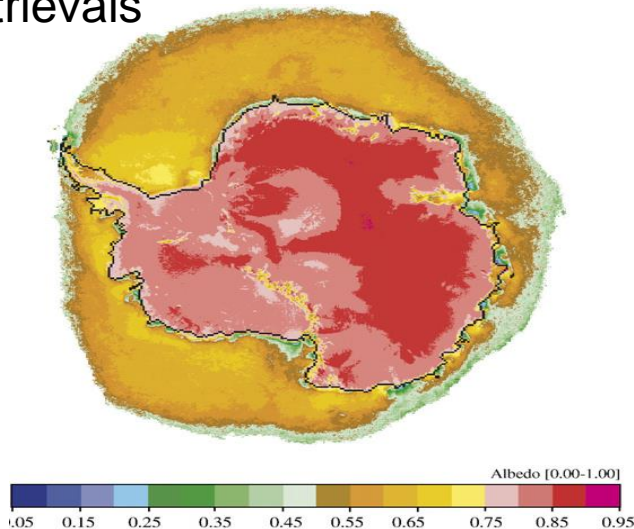
Antarctic ice sheet and sea ice regional albedo and temperature change,  
1981–2000, from AVHRR Polar Pathfinder data

Vesa Laine

*Finnish Meteorological Institute, PO Box 503, 00101 Helsinki, Finland*

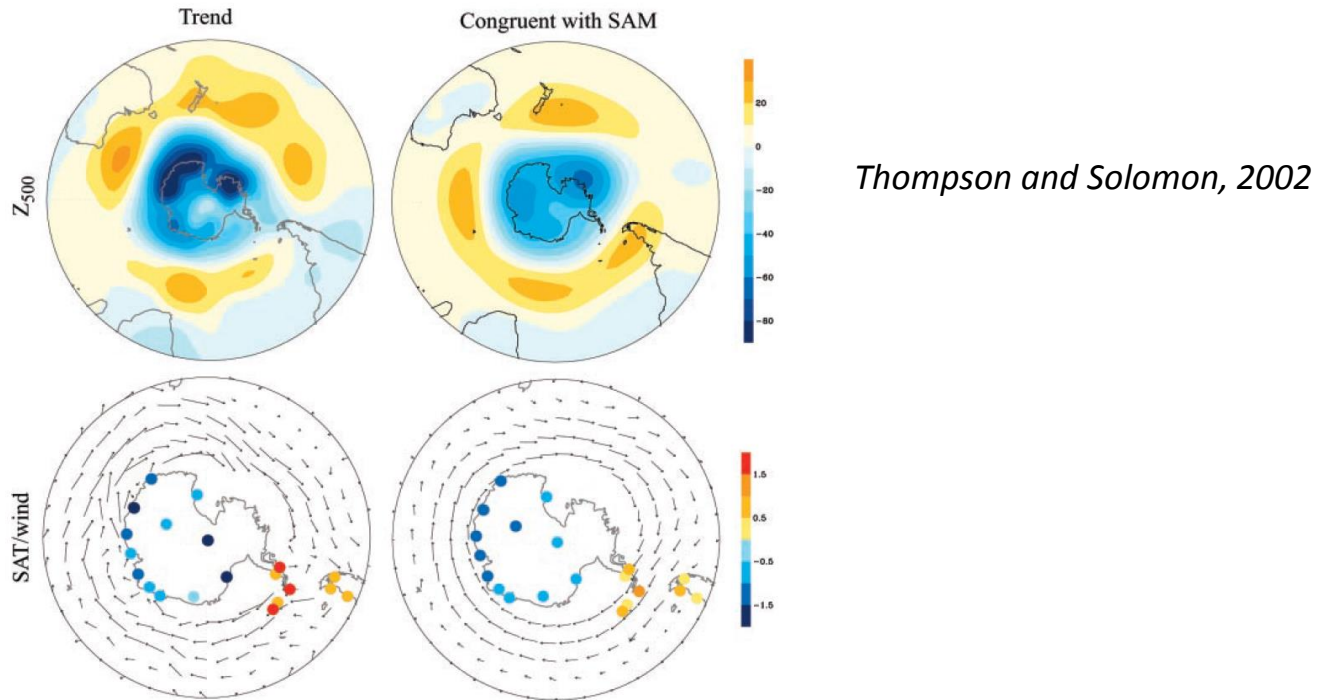
Received 9 January 2007; received in revised form 1 June 2007; accepted 2 June 2007

Cloud masked albedo  
retrievals



... positive trend for ice sheet over 1980-2000

# Did ozone depletion have an impact on Antarctic climate?



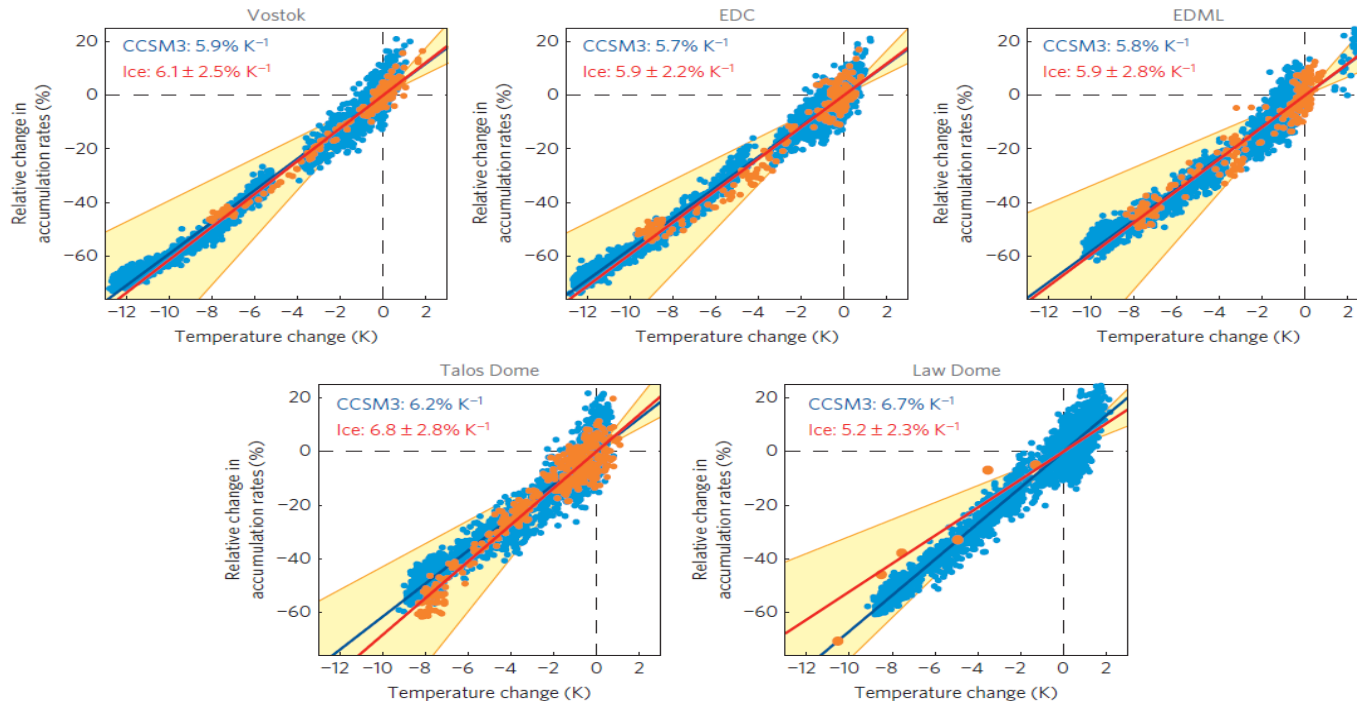
Cooling over E-Antarctica, and warming of Peninsula  
(SAM-driven pattern, ...so “indirectly” induced.)

# Snowfall increase over Antarctica

## Increase in accumulation with warming

### Consistent evidence of increasing Antarctic accumulation with warming

Katja Frieler<sup>1\*</sup>, Peter U. Clark<sup>2</sup>, Feng He<sup>2,3</sup>, Christo Buizert<sup>2</sup>, Ronja Reese<sup>1,4</sup>, Stefan R. M. Ligtenberg<sup>5</sup>, Michiel R. van den Broeke<sup>5</sup>, Ricarda Winkelmann<sup>1,4</sup> and Anders Levermann<sup>1,4</sup>



**Figure 1 | Changes in local accumulation rates and temperatures derived from ice cores (orange) and CCSM3 palaeo-simulations (blue, decadal averages) at the ice-core sites.** Changes in accumulation and temperature are described in comparison to a core-specific pre-industrial reference level (see Supplementary Information). Thick solid lines are derived by linear regression assuming that the intercept is zero (orange lines for ice-core data and blue lines for simulations, sensitivities are given in each panel including the  $2\sigma$  uncertainty range of the sensitivities derived from the ice cores). The shaded area describes the uncertainty range of the ice-core sensitivities.

# Negative snowfall–albedo feedback?

## Can snowfall reduce the positive snow-albedo feedback?

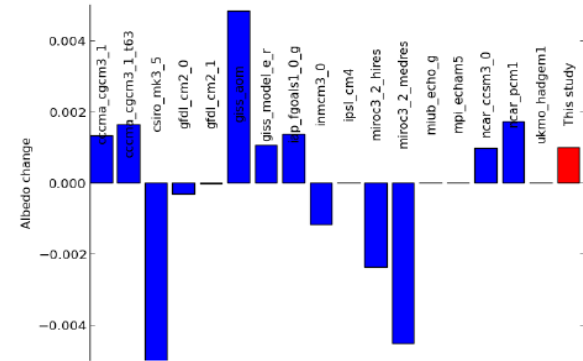


### Inhibition of the positive snow-albedo feedback by precipitation in interior Antarctica

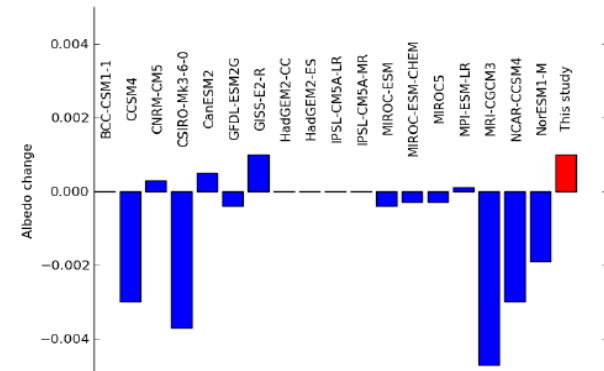
G. Picard<sup>1\*</sup>, F. Domine<sup>1,2</sup>, G. Krinner<sup>1</sup>, L. Arnaud<sup>1</sup> and E. Lefebvre<sup>1</sup>

- Projected future **increase** in **precipitation** can **increase** snow **albedo** by 0.4% on average over 21st century and **overcompensate** expected **albedo decrease** owing to **warming**

- compensation seems model dependent

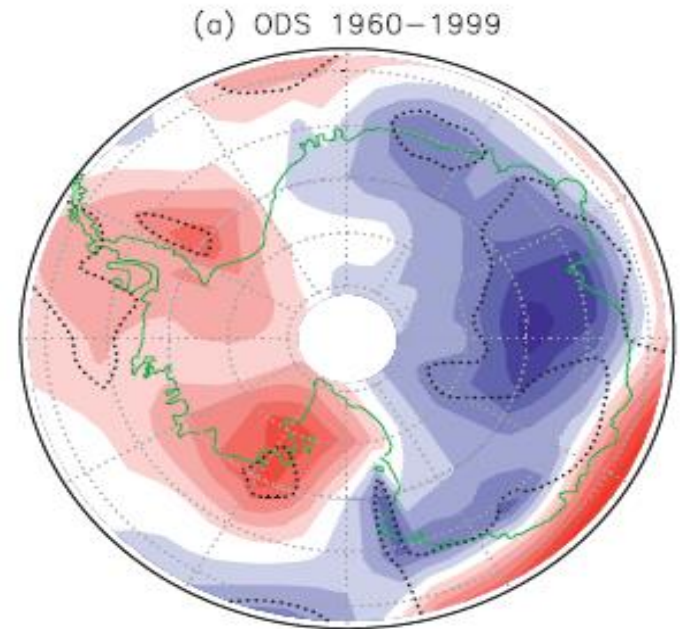
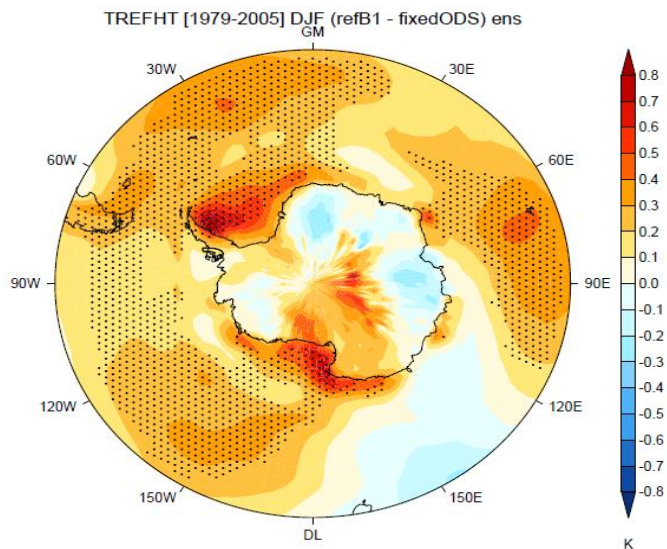
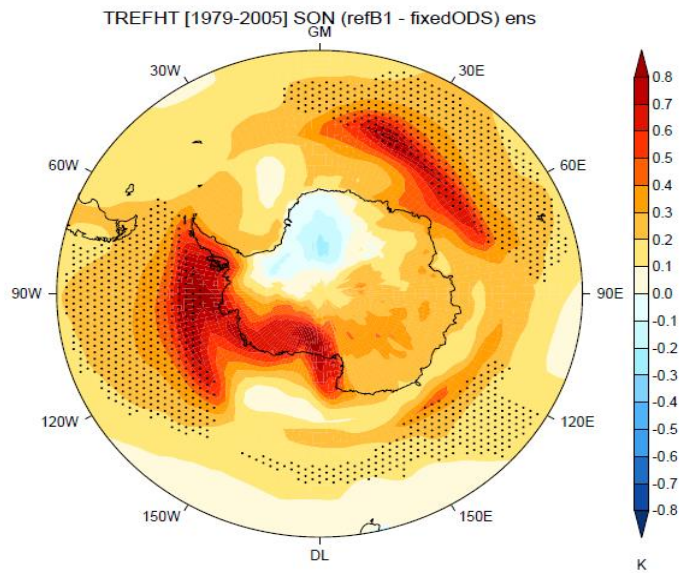


b)

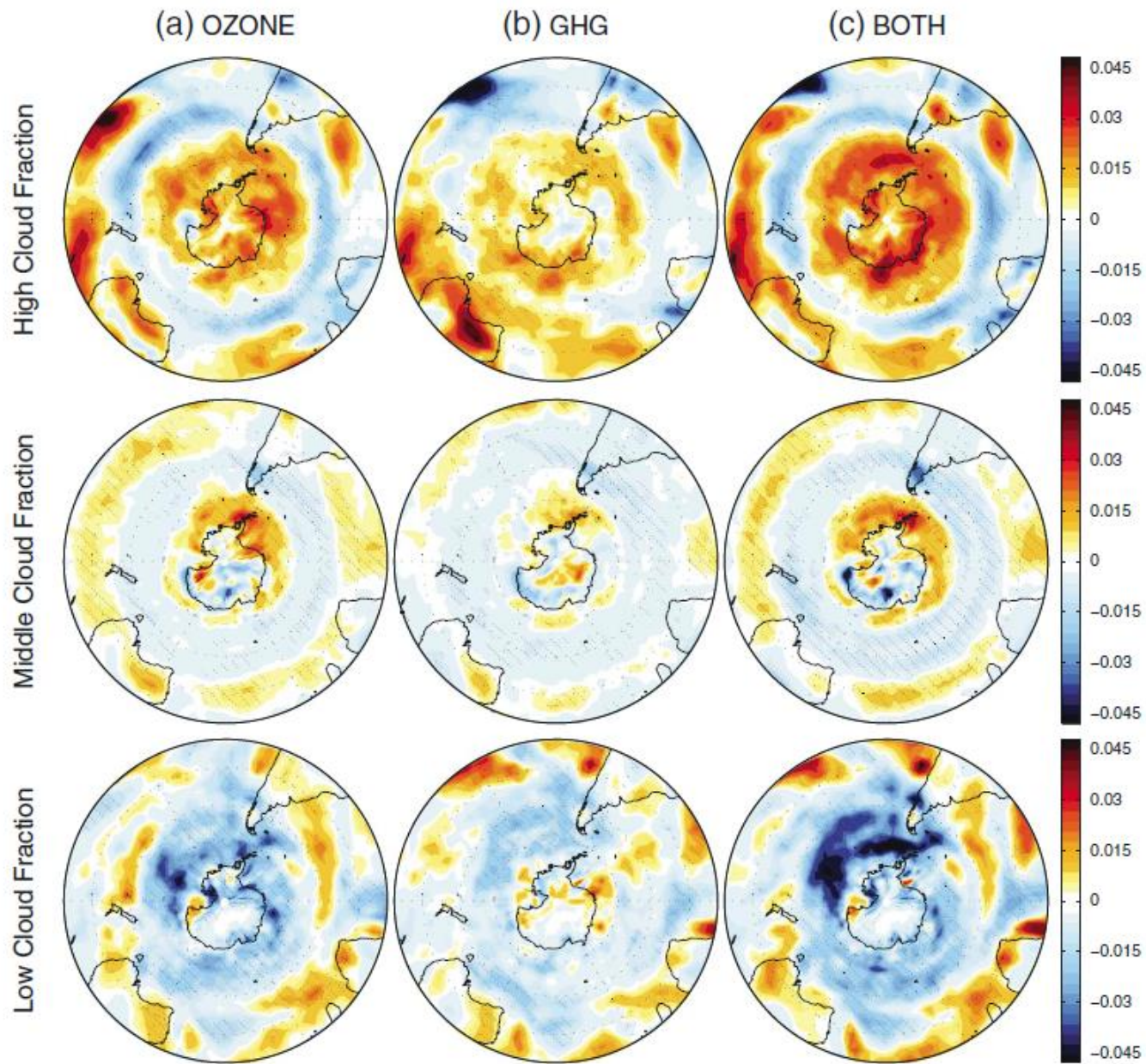


Supplementary Figure 5. Albedo change in Antarctica (above 2000m) predicted by models in the Coupled Model Intercomparison Project a) CMIP3 under SRES A1B scenario and b) CMIP5 under RCP4.5 scenario between the present (1986-2005) and future (2080-2099). The increase of albedo estimated in this study at Dome C is shown in red.

# Does O3-hole “cool” E-Antarctica?



*McLandress et al., 2011*



*Grise et al., 2013*